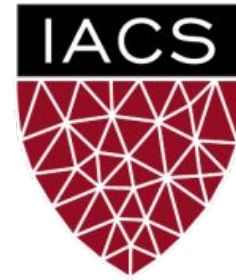


An Information Theory Approach on Deciding Spectroscopic Follow Ups



JAVIERA ASTUDILLO
PAVLOS PROTOPAPAS
KARIM PICCHARA
PABLO HUIJSE

OUTLINE



1 LSST

2 VARIABLE OBJECTS

3 TYPES OF DATA

4 BACKGROUND THEORY

5 RELATED WORK

6 PROBLEM DESCRIPTION

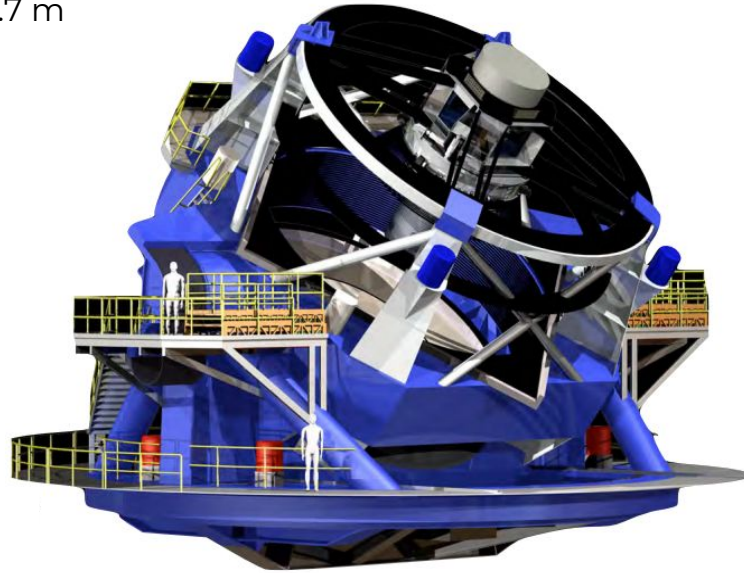
7 METHODOLOGY

8 RESULTS

9 CONCLUSION AND FURTHER WORK

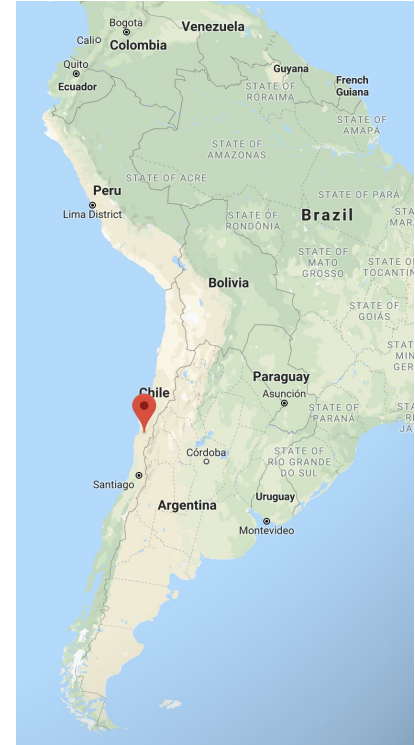
LSST

- Effective aperture of 6.7 m
- FoV 9.6 deg²
- Large *etendue*
(Collecting area x FoV)



Wide-Deep-Fast
Cover large swaths of sky to
faint magnitudes in a short
amount of time.

2022-2032



BACK OF THE ENVELOPE LSST

Sky area: 41252.96 deg², LSST field of view (FoV): 9.6 deg²
→ southern sky area ~ 20000 deg² ~ 2000 LSST FoV

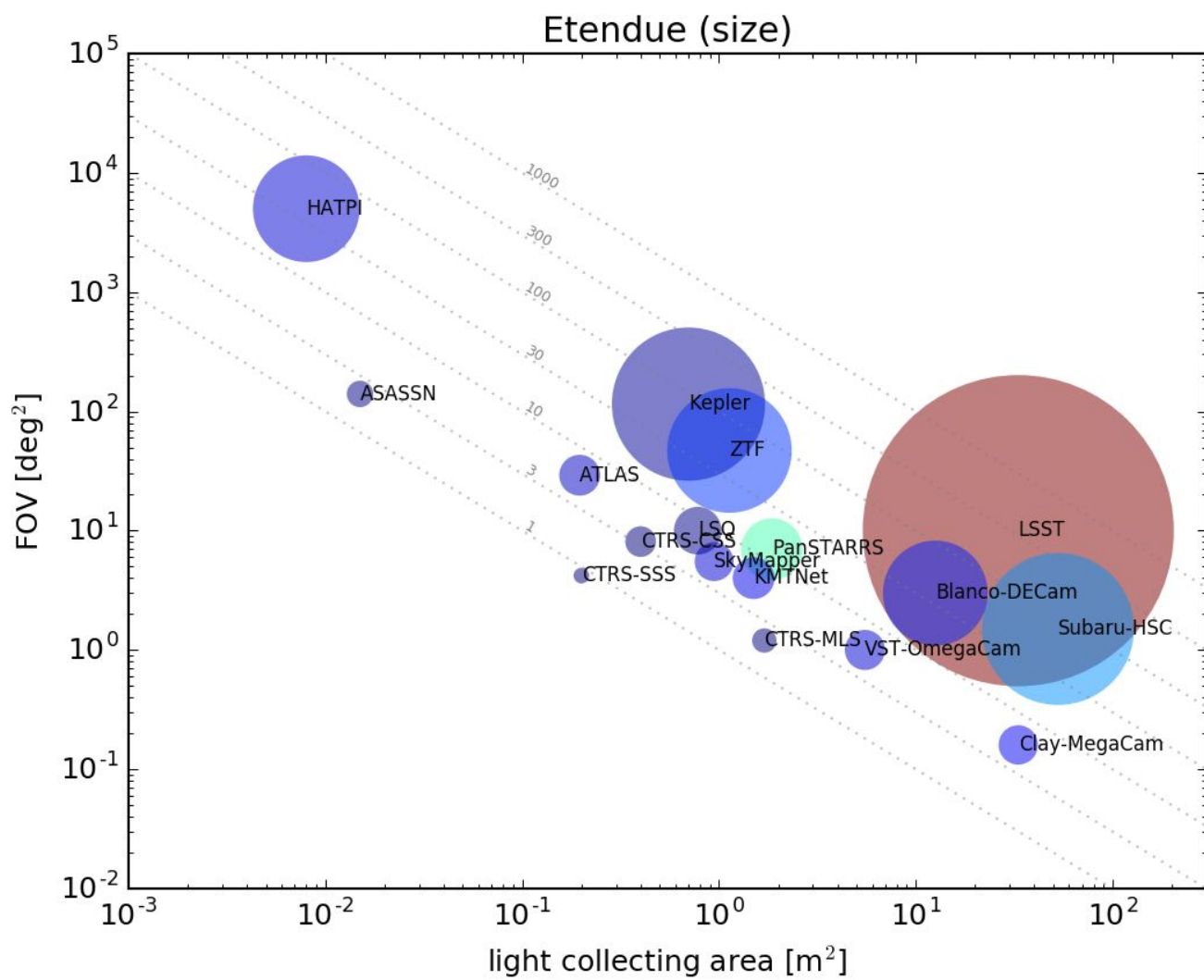
Night length ~ 8 hr, LSST exposure: 30 + 7 sec
→ ~800 LSST exposures / night

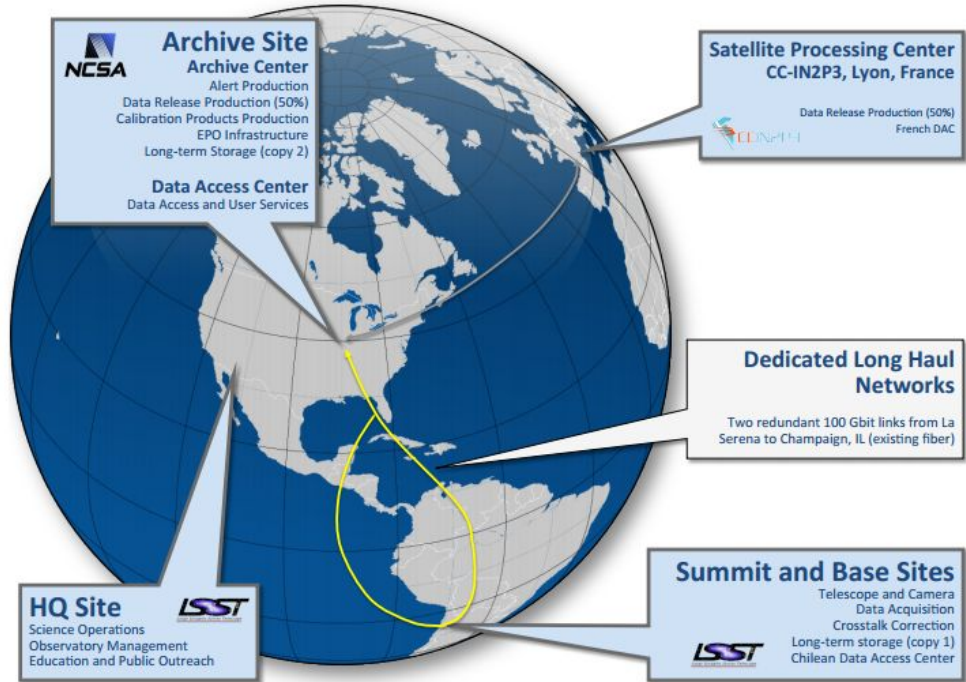
Survey length ~ 10 yr (-10%) in 6 bands (ugrizy)
→ 365 x 9 x 800 ~ 2.5M LSST exposures
→ ~2.5M visits / 2000 FoV = 1250 visits per FoV
→ average of one visit every 3 days per FoV
→ ~250 visits per region per band
→ average of one visit every 15 days per FoV per band

Single epoch depth in r band ~ 24.5 mag
→ Coadded depth of 24.5 + 2.5 log₁₀(sqrt(250)) ~ 27.5

LSST's goal is to have 1000 visits to every region of the southern sky.

Exact distribution of visits TBD: different science goals can have complementary/conflicting optimal strategies (e.g. proper motion studies vs supernova studies, a.k.a. cadence wars).





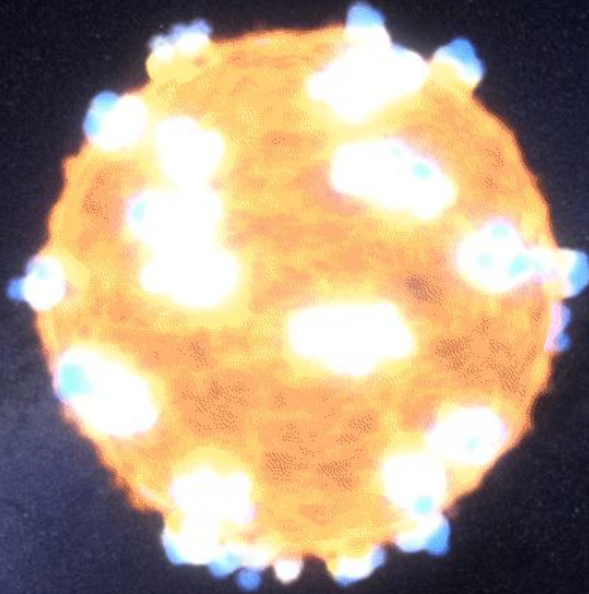
1



2

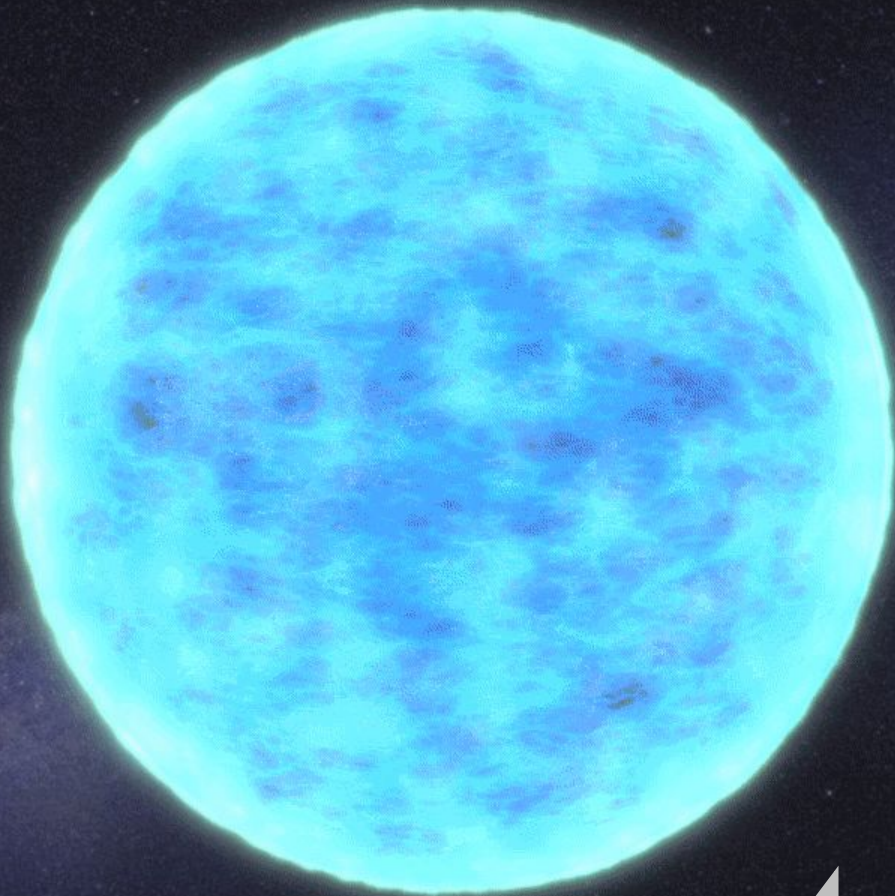
VARIABLE OBJECTS

2



SOURCE. NASA Ames, STScI/G. Bacon

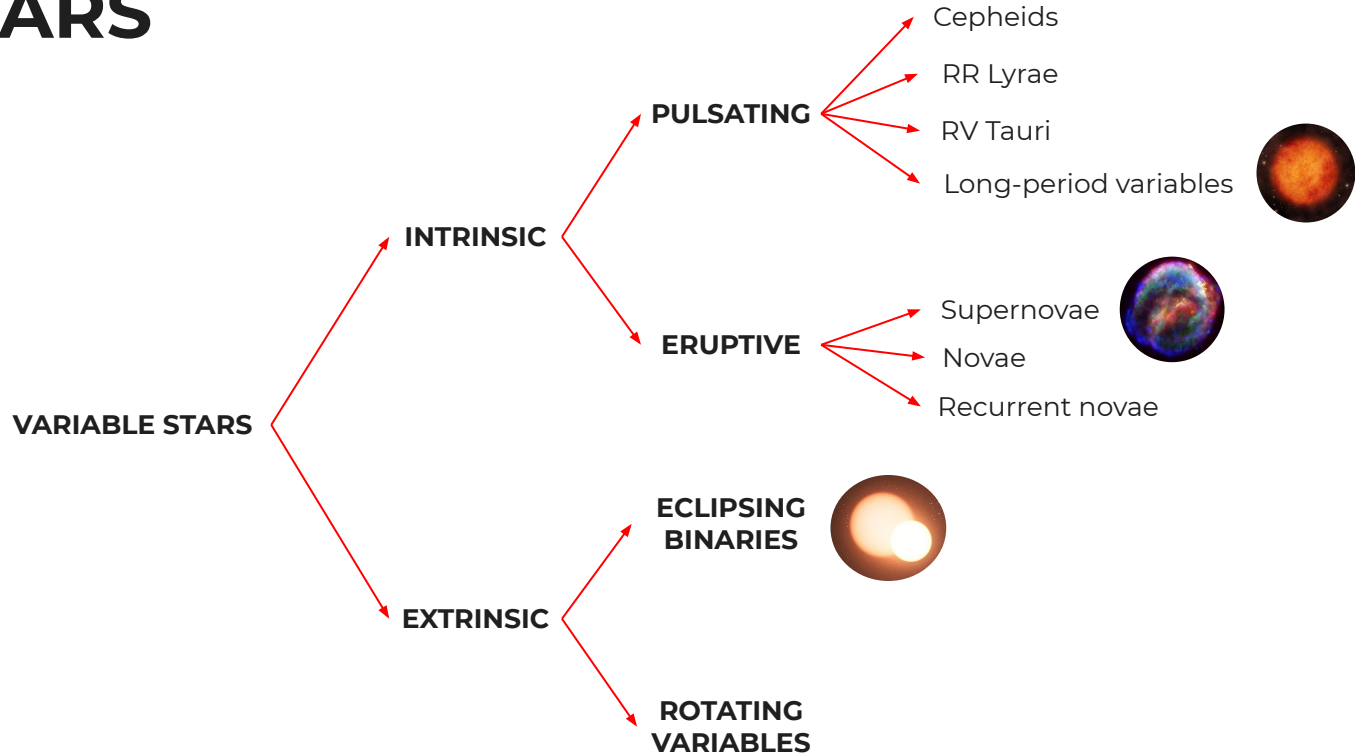
2



SOURCE. NASA Ames, STScI/G. Bacon

VARIABLE STARS

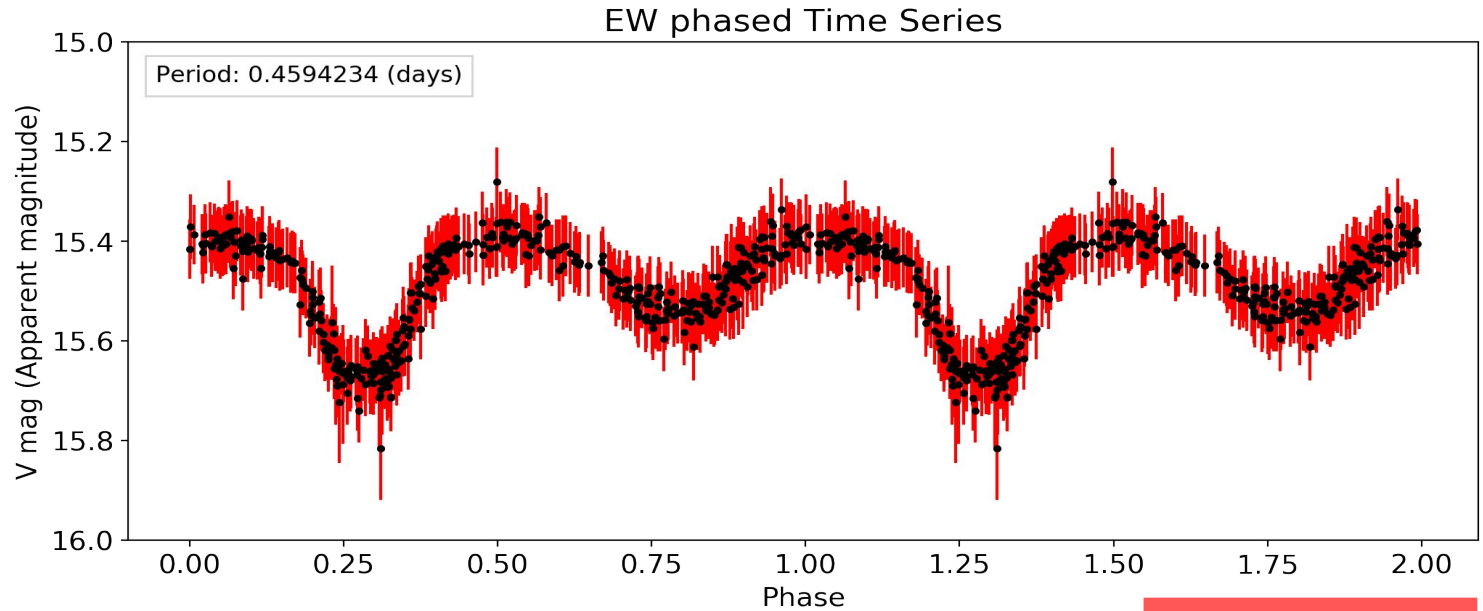
2



TYPES OF DATA

TIME SERIES

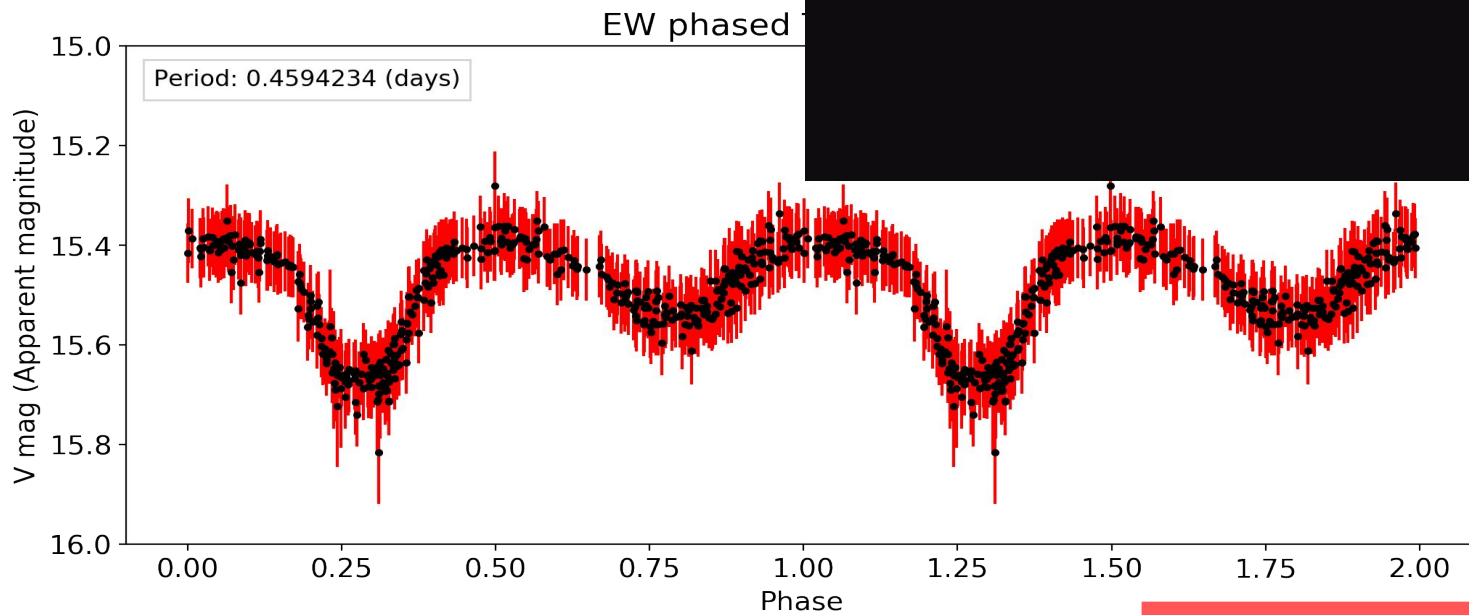
3



SOURCE. CSS DRI

TIME SERIES

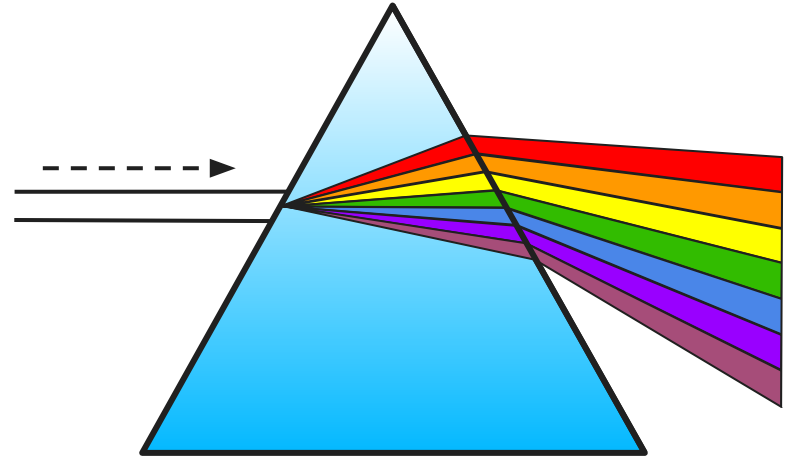
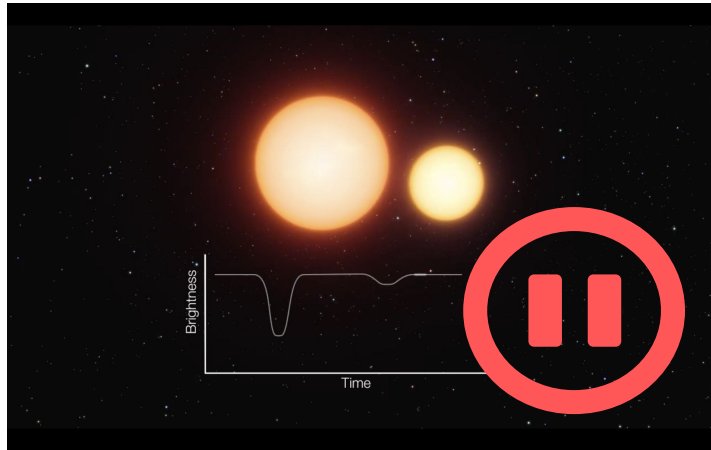
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SOURCE. CSS DR1

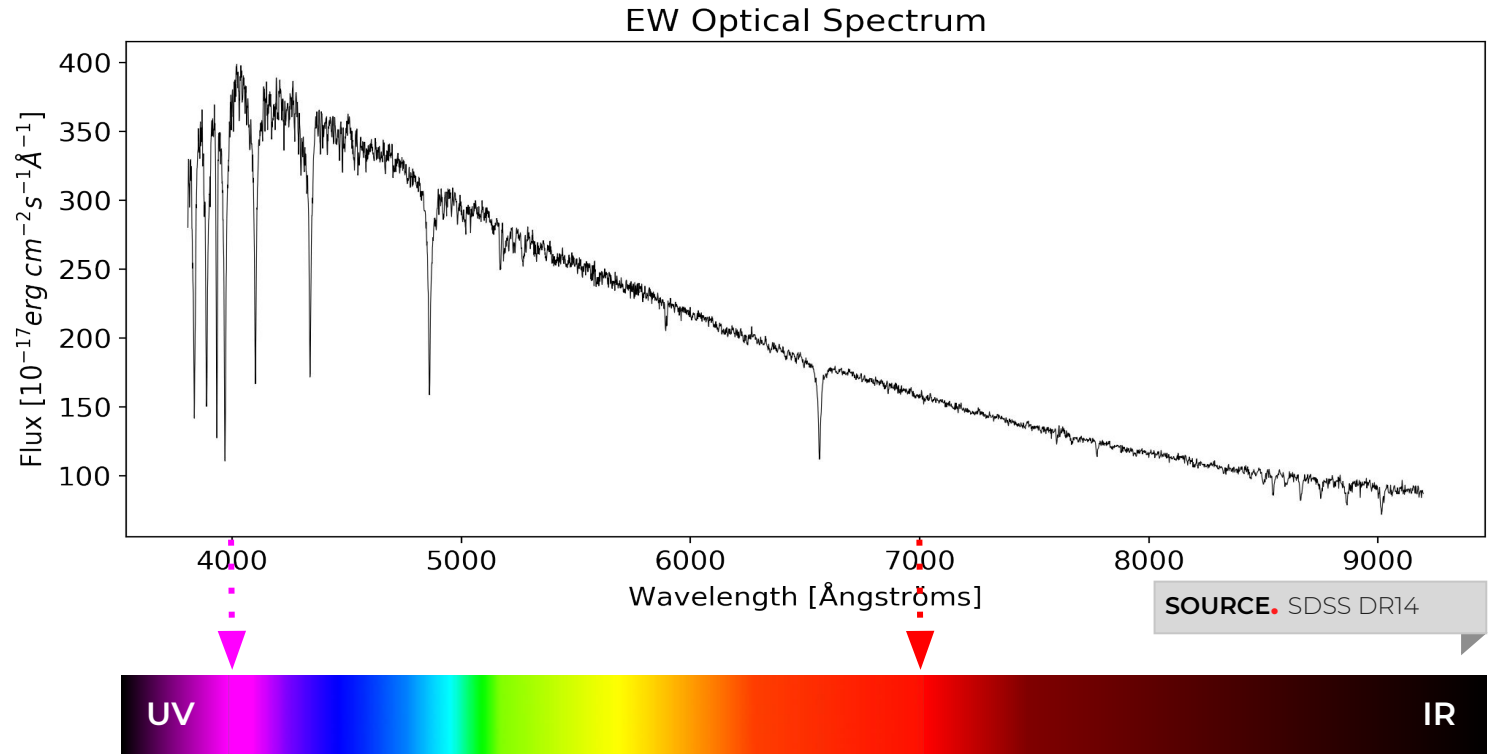
SPECTRUM

3



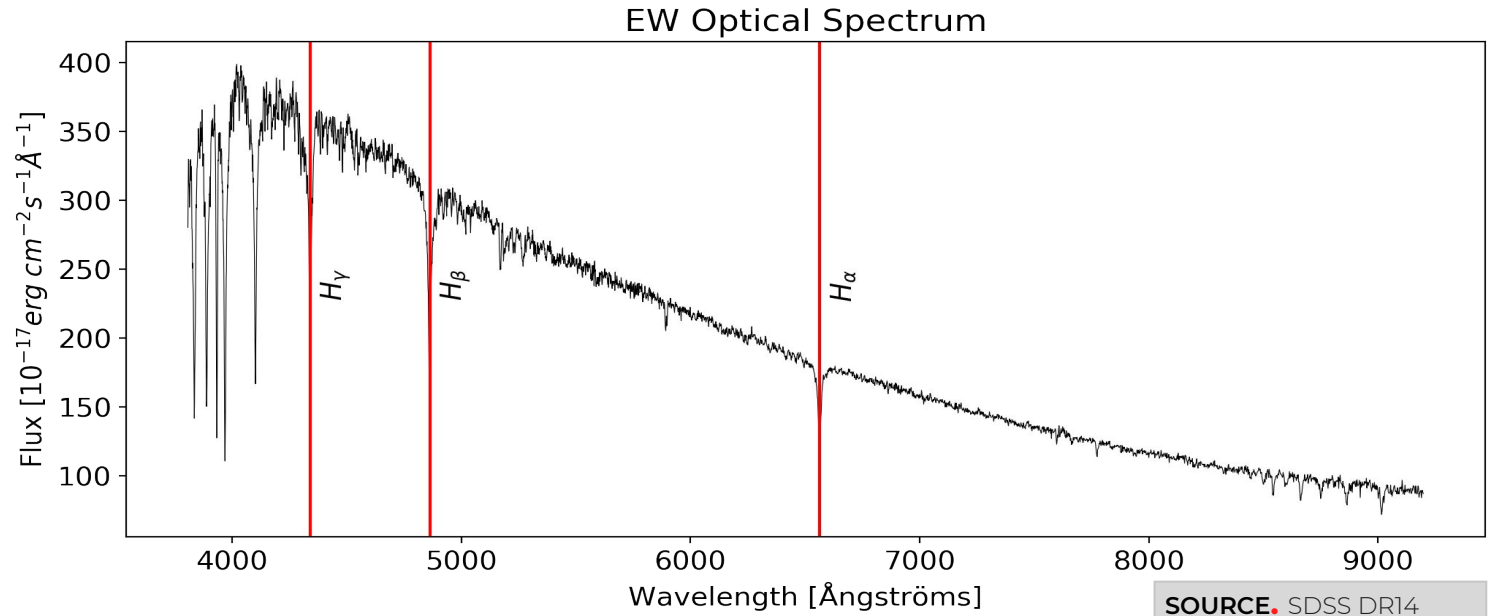
SPECTRUM

3



SPECTRUM

3



DATA COLLECTION



TIME SERIES

30 s./Exposure
Many

4 EXPOSURES
200 OBSERVATIONS
400 MIN

CSS DR1



SPECTRUM

15 m./Exposure
Few

3 EXPOSURES
1 OBSERVATION
45 MINUTES

SDSS DR14



3


IDEAL SCENARIO

TIME SERIES



DATASET
100%

SPECTRUM



DATASET
100%


REALITY

TIME SERIES



DATASET
100%

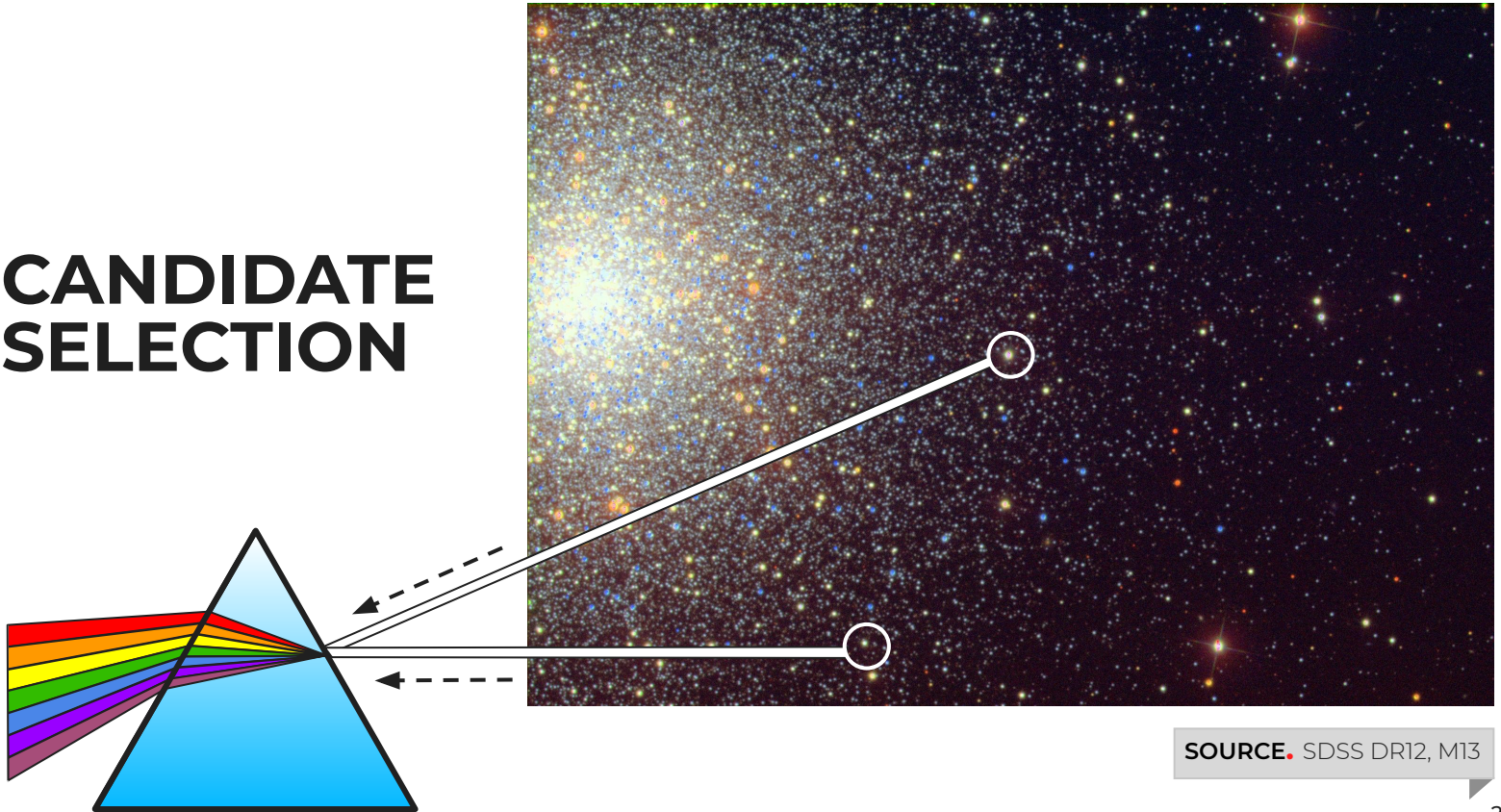
SPECTRUM



DATASET
<<100%

3

CANDIDATE SELECTION



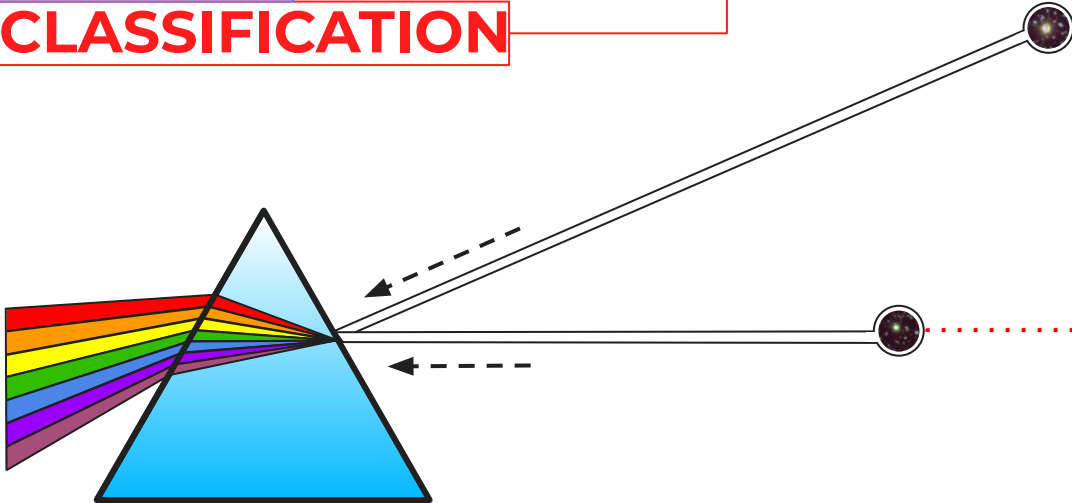
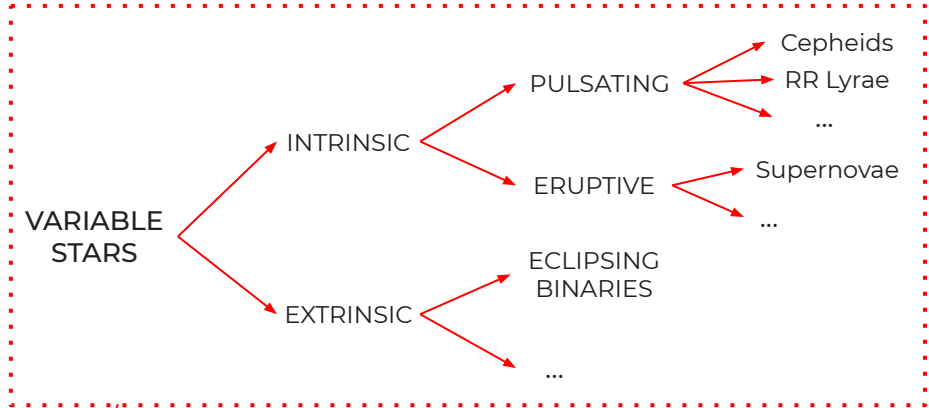
SOURCE. SDSS DR12, M13

3

+ HIGHER ACCURACY

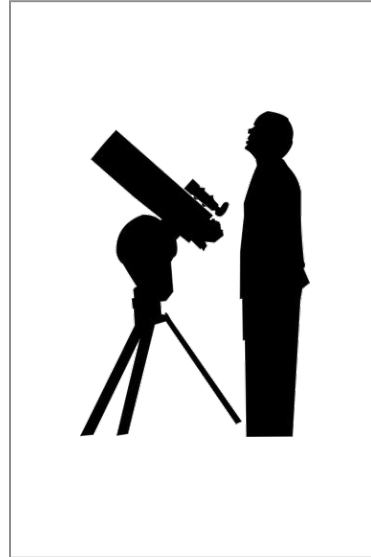
+ HIGHER CONFIDENCE IN THE CLASSIFICATION

IMPROVE CLASSIFICATION



3

IMPROVE CLASSIFICATION



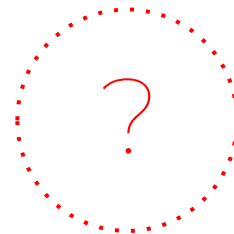
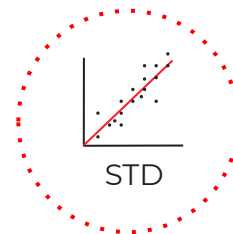
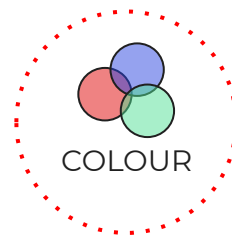
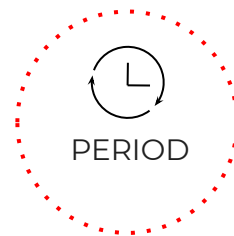
BACKGROUND THEORY

CLASSIFICATION



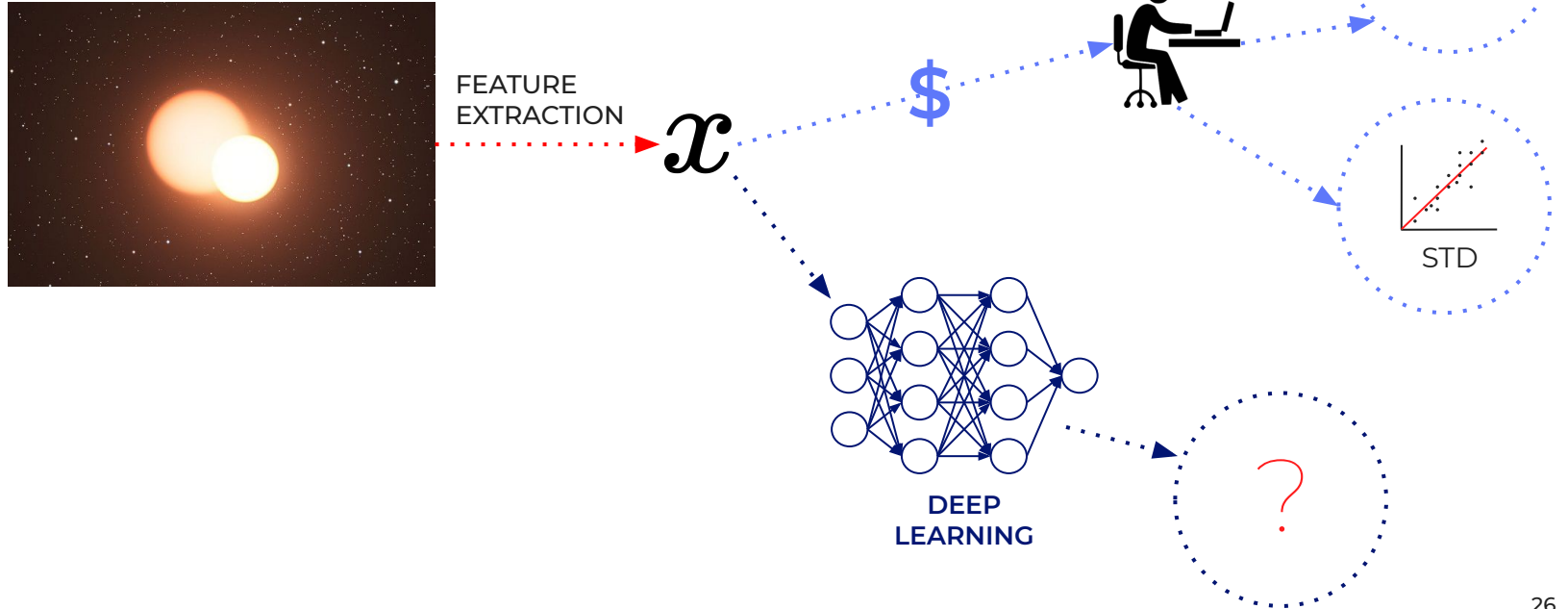
FEATURE EXTRACTION

x



CLASSIFICATION

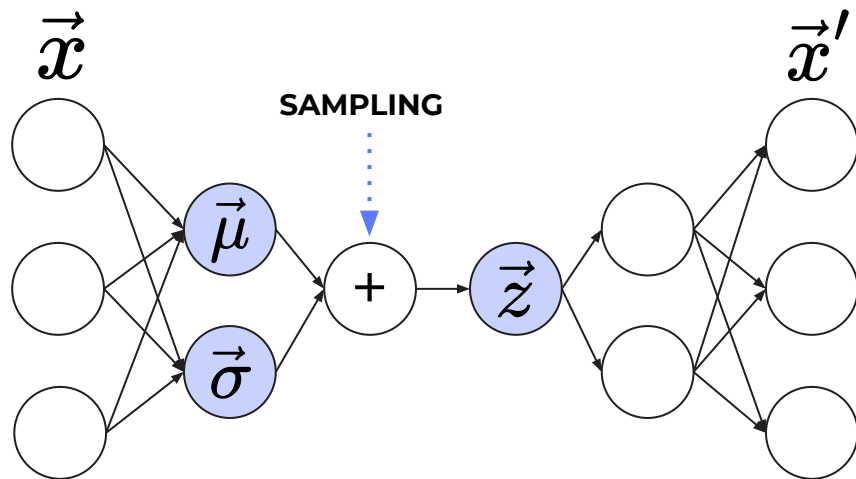
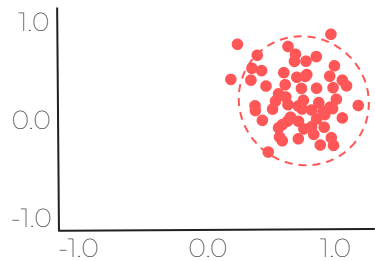
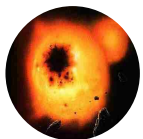
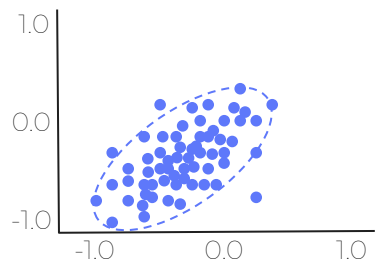
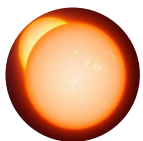
4



4

FEATURE EXTRACTION

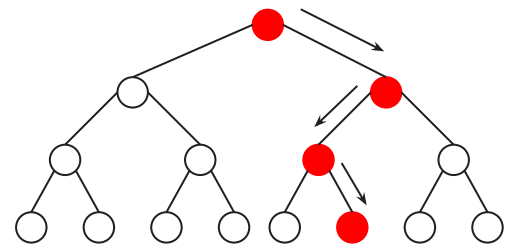
\vec{z} : ENCODING SPACE



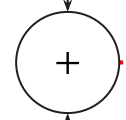
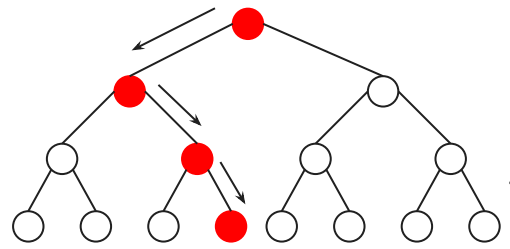
4

x →

RANDOM FOREST



...

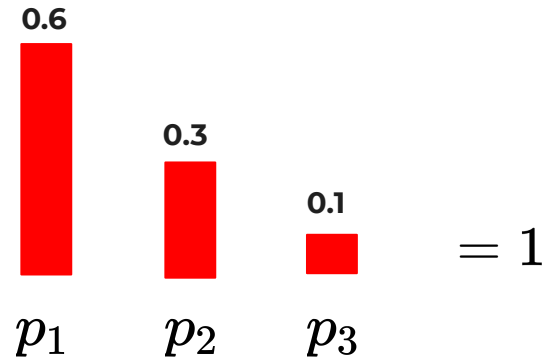


CLASSIFY

→ $P(\hat{y} = c|x)$

4

ENTROPY

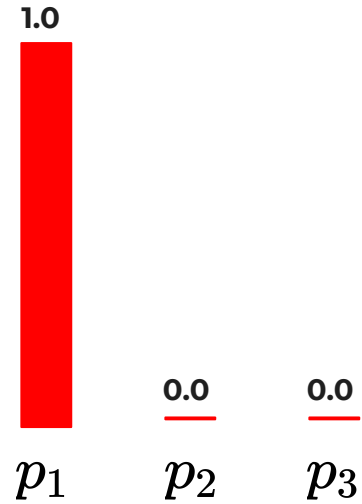


$$H = - \sum_{i=1}^n p_i \log p_i$$

$$H = 0.389$$

4

ENTROPY

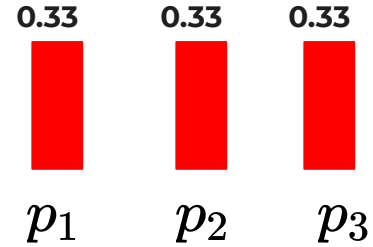


LEAST CONFUSED!

$$H = 0$$

4

ENTROPY

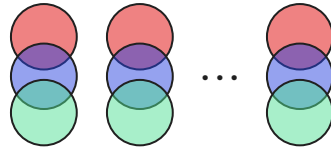


MOST CONFUSED!

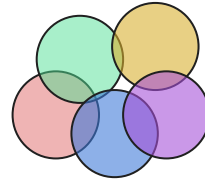
$$H = 0.477$$

RELATED WORK

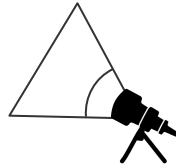
RELATED WORK



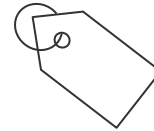
PETERS ET AL.
(2015)



YANG
(2015)



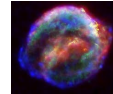
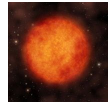
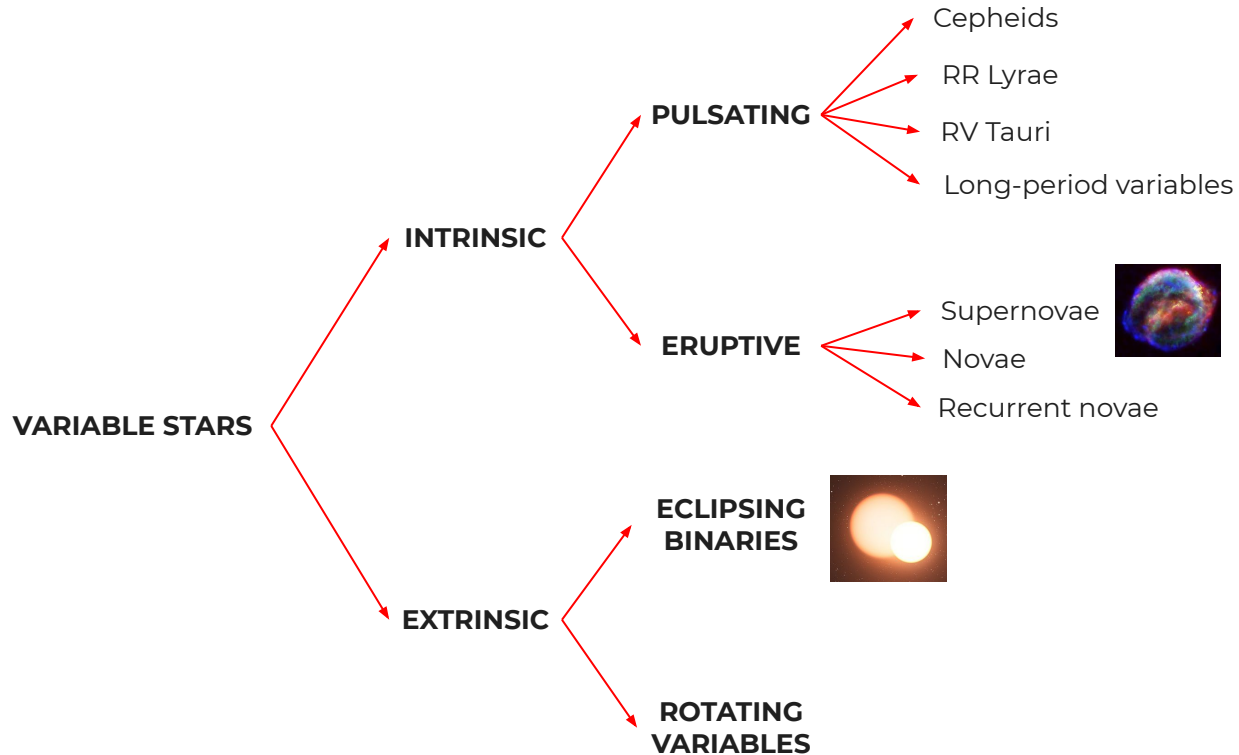
XIA
(2016)



ISHIDA
(2019)

PROBLEM DESCRIPTION

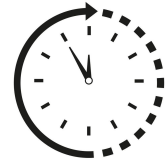
CLASSIFICATION



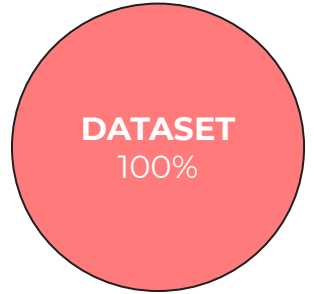
6



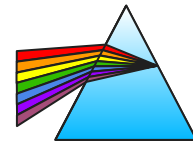
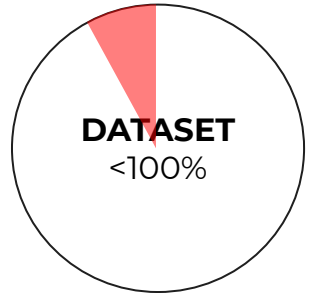
FEATURE EXTRACTION



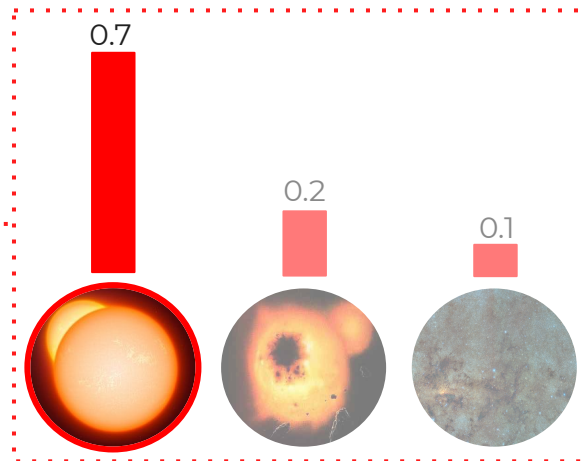
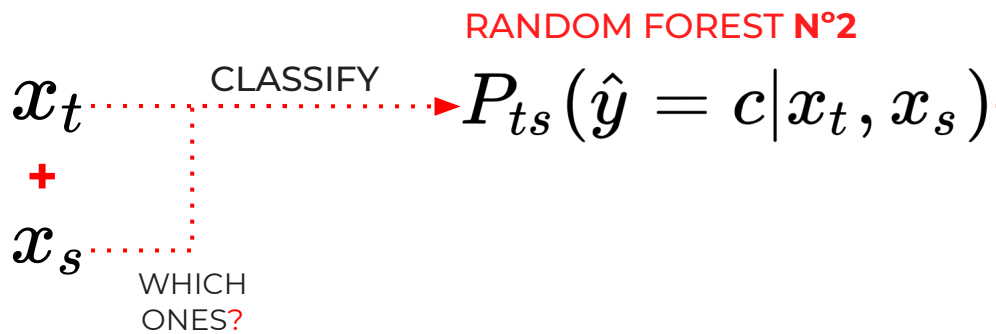
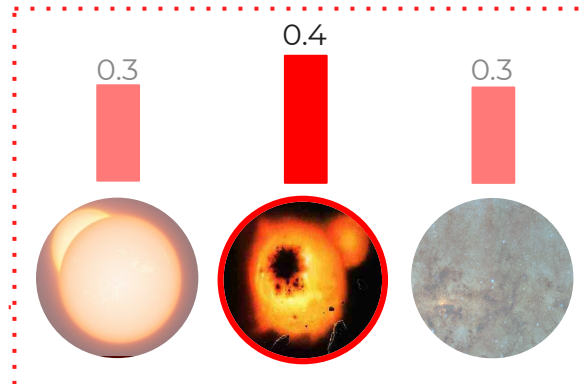
\mathcal{X}_t



\mathcal{X}_s



6



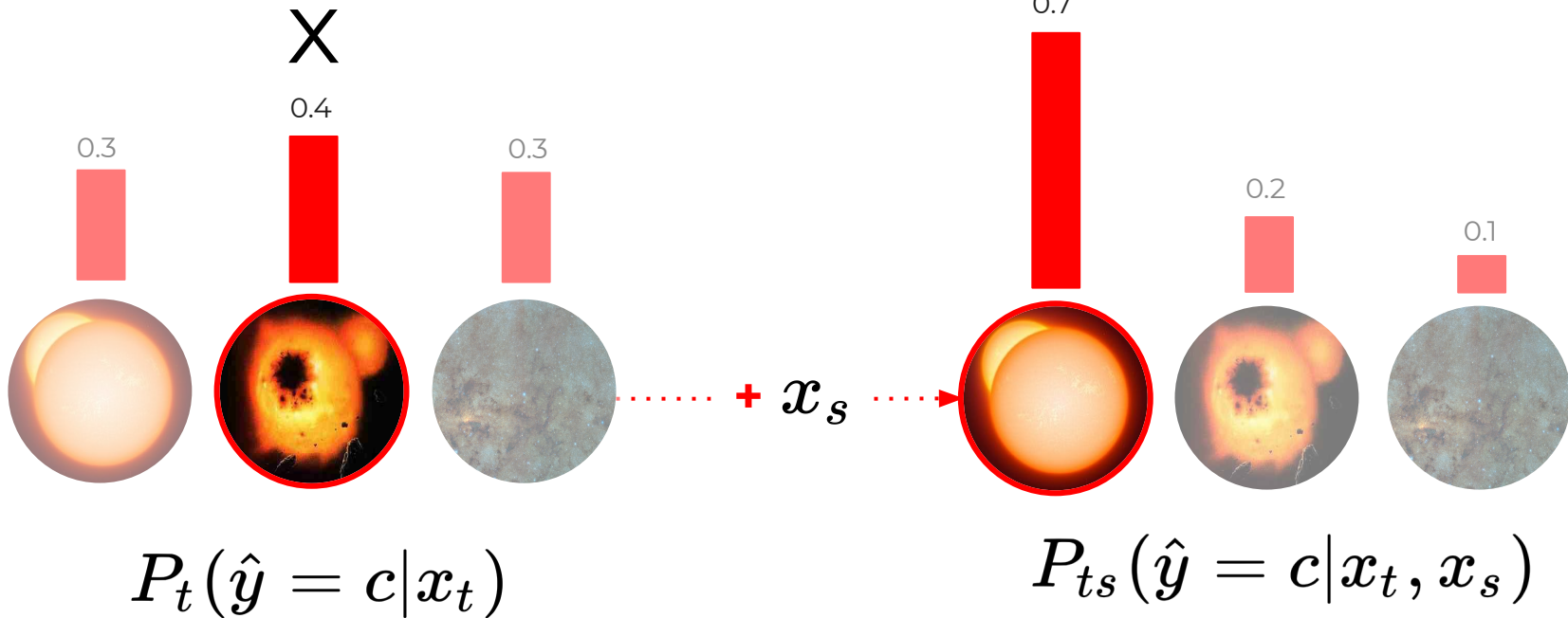
6

WHICH ONES?

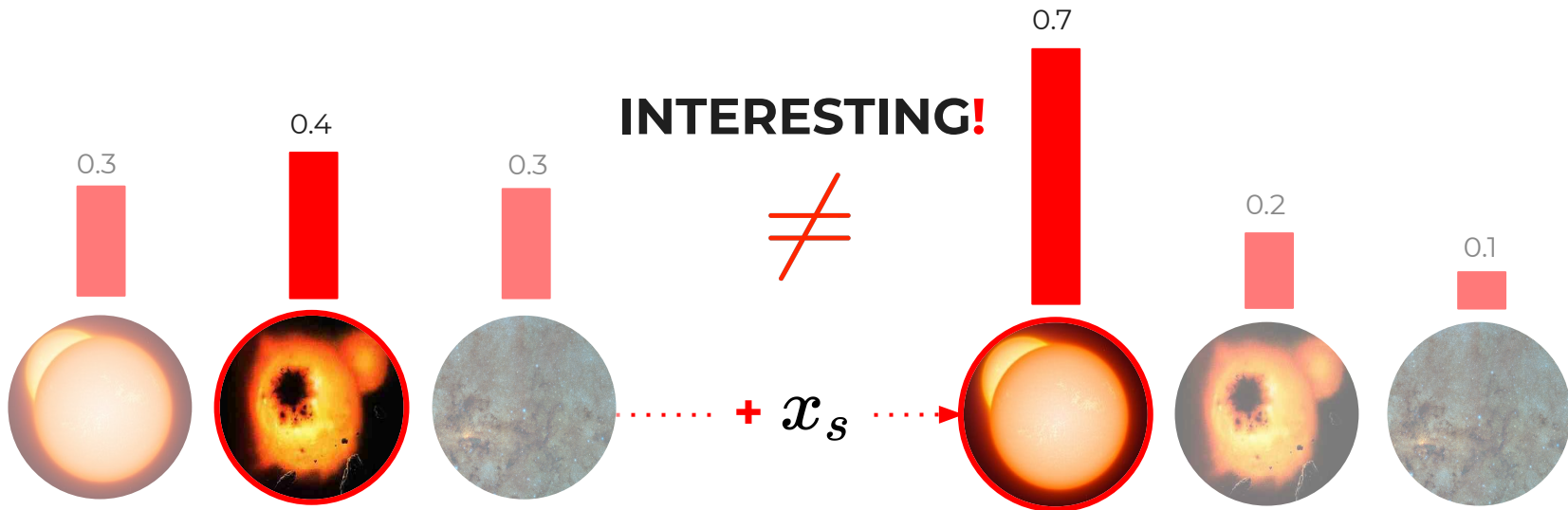
CASE 1

ACCURACY

6



6

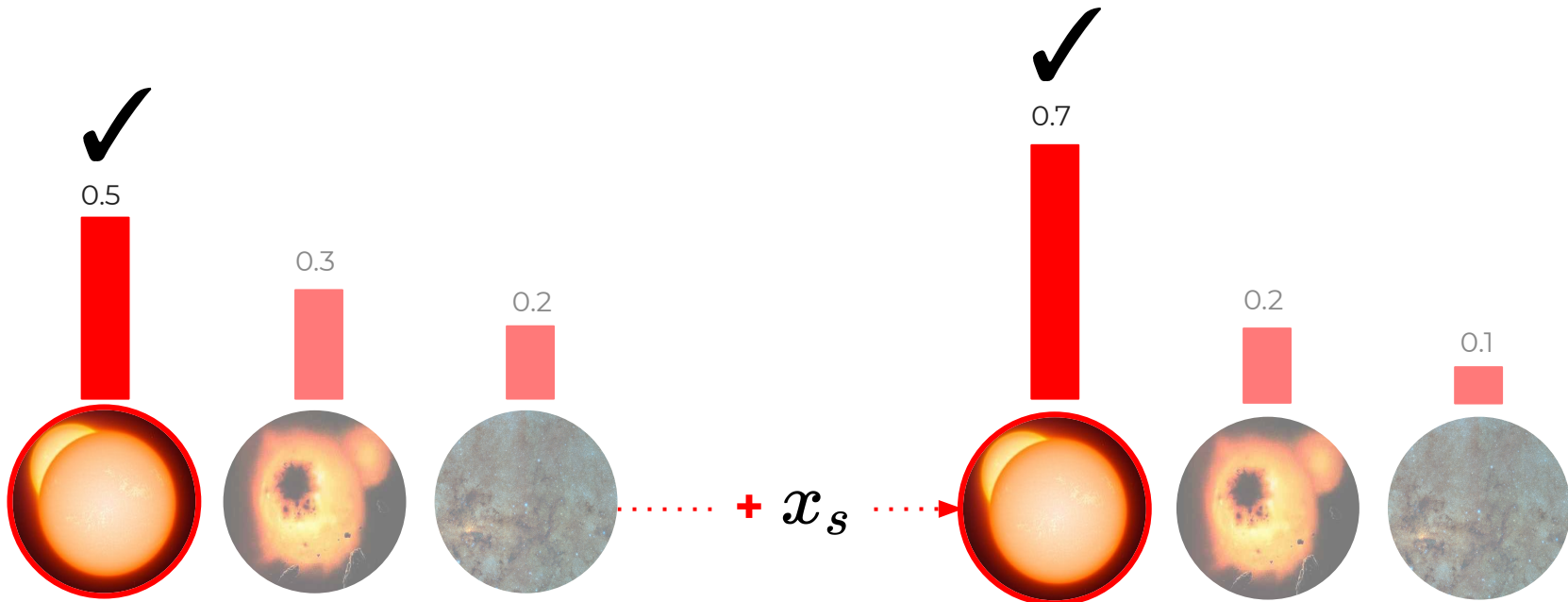


$$\Delta \hat{y}(x_t, x_s) = 1$$

6

WHICH ONES?
CASE 2
CONFIDENCE

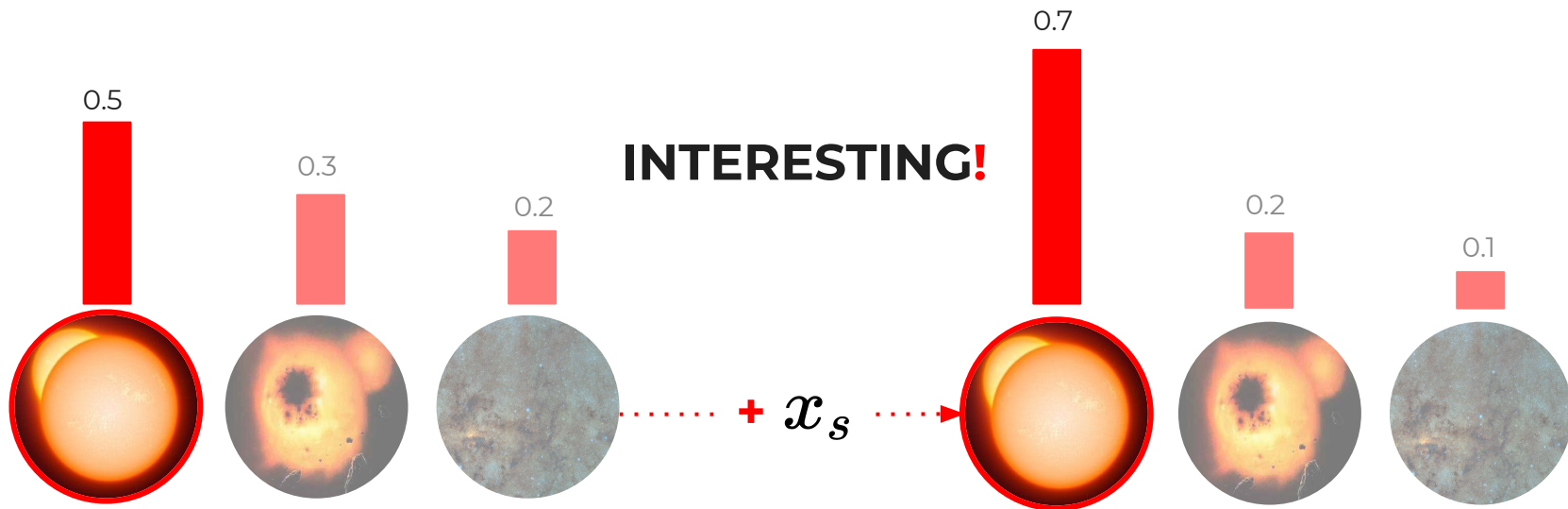
6



$$P_t(\hat{y} = c | x_t)$$

$$P_{ts}(\hat{y} = c | x_t, x_s)$$

6



$$H_t(\hat{y}|x_t)$$

$$>$$

$$H_{ts}(\hat{y}|x_t, x_s)$$

$$IG(x_t, x_s) = H_t(\hat{y}|x_t) - H_{ts}(\hat{y}|x_t, x_s) > 0$$

METRICS

CASE 1

ACCURACY
Formerly wrongly
classified

$$\Delta \hat{y}(x_t, x_s) = 1$$

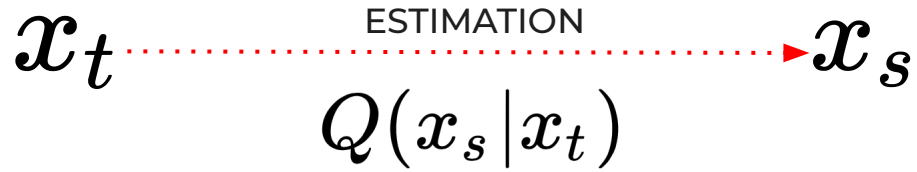
CASE 2

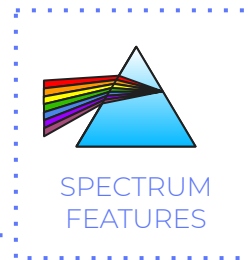
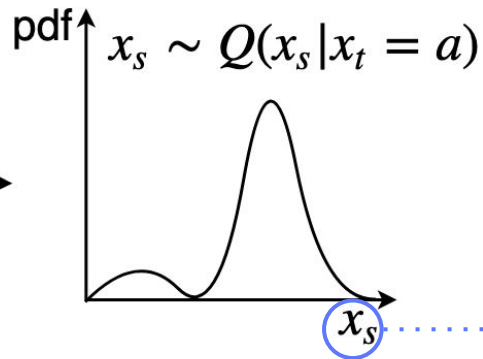
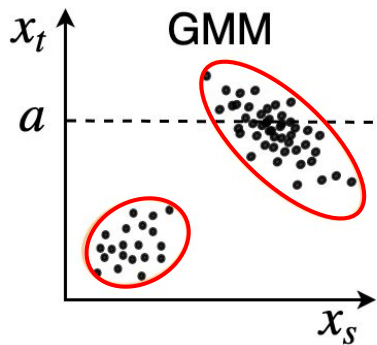
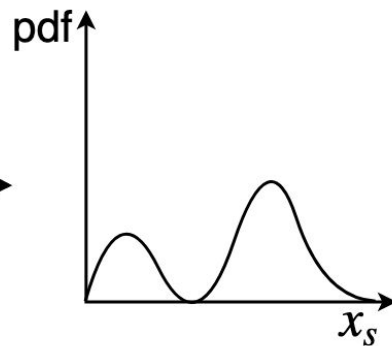
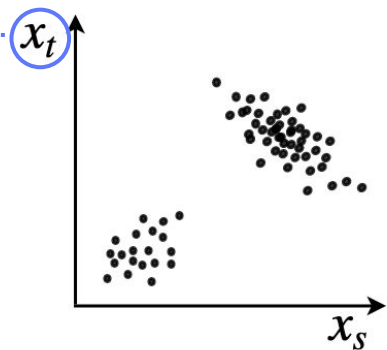
CONFIDENCE
Less confused
outcome

$$IG(x_t, x_s) > 0$$

METHODOLOGY

7

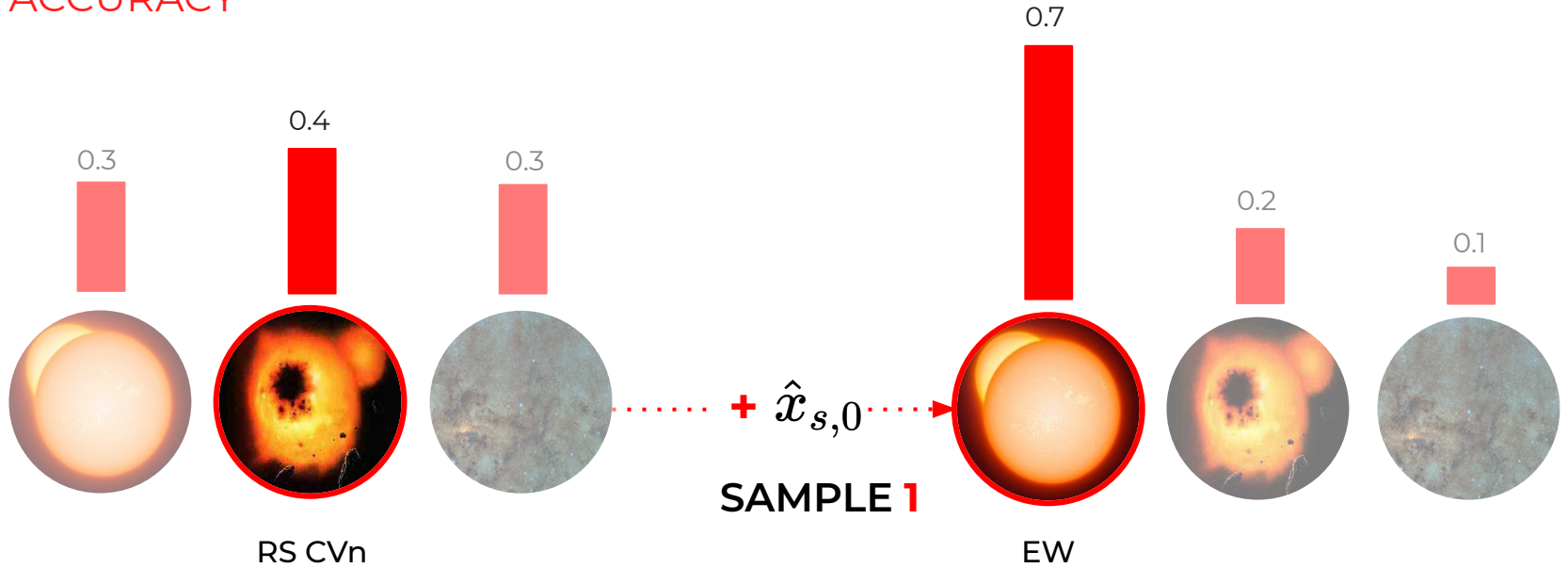




CASE 1

ACCURACY

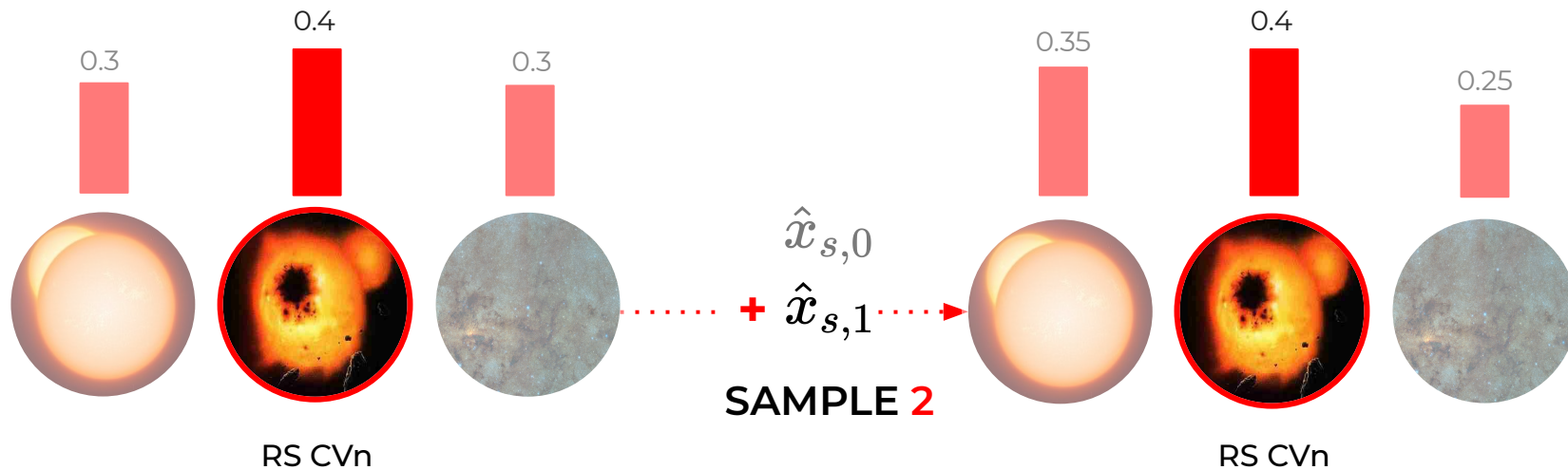
7



CASE 1

ACCURACY

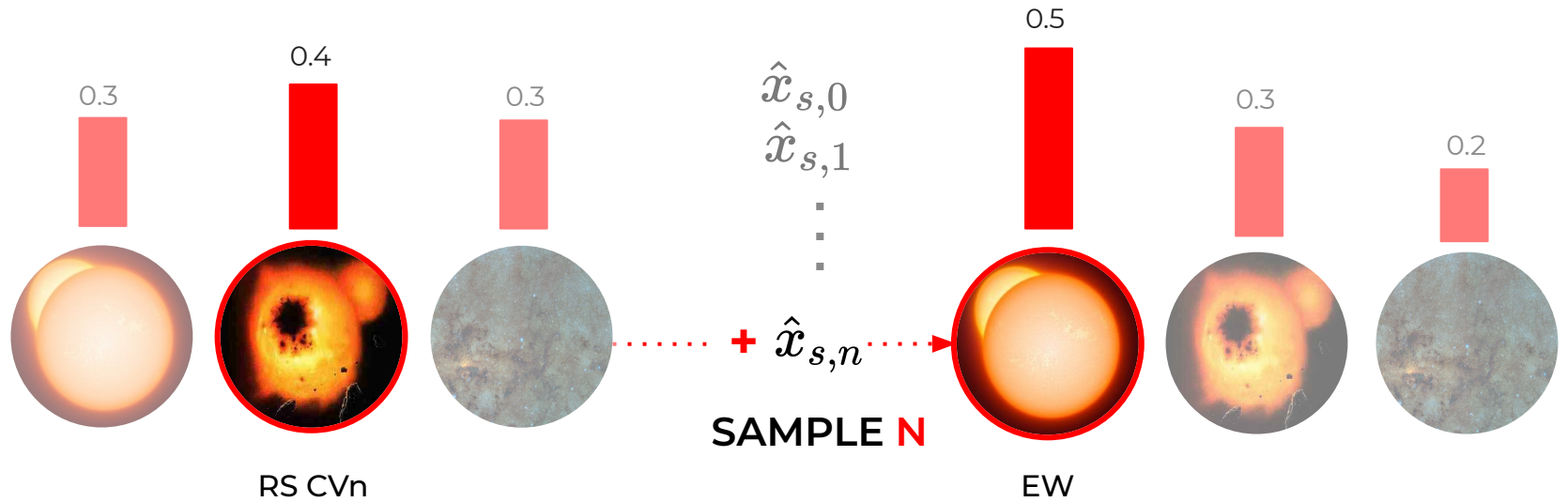
7



CASE 1

ACCURACY

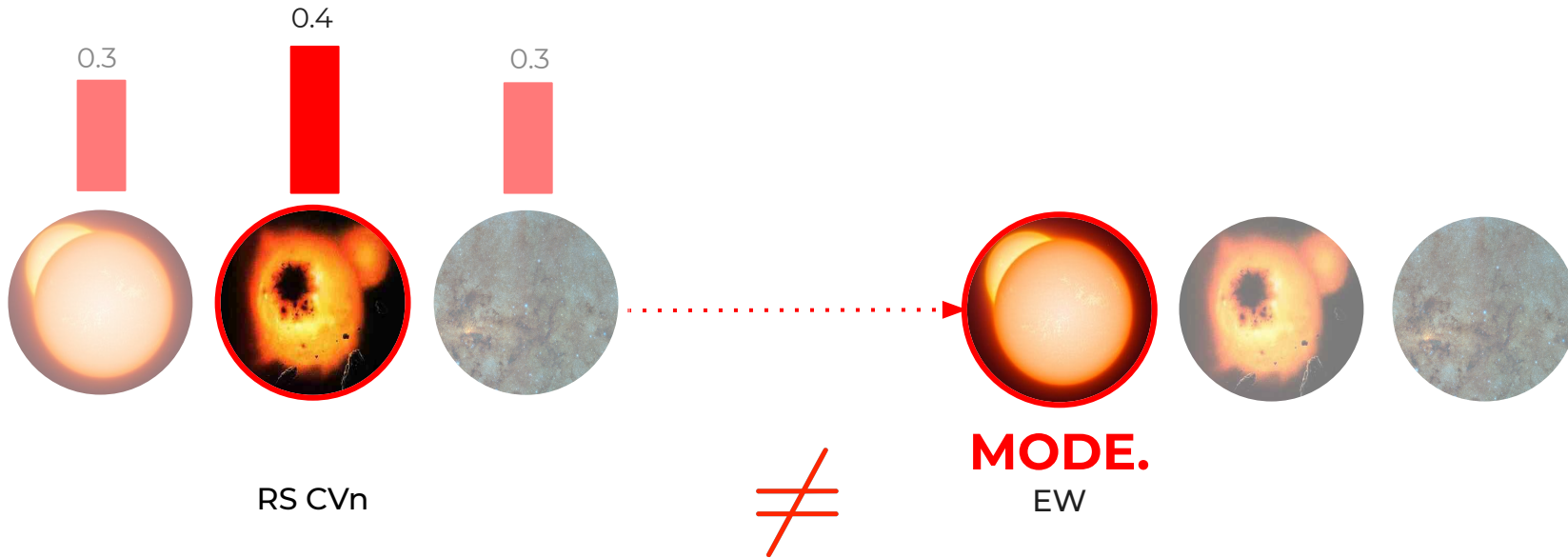
7



CASE 1

ACCURACY

7



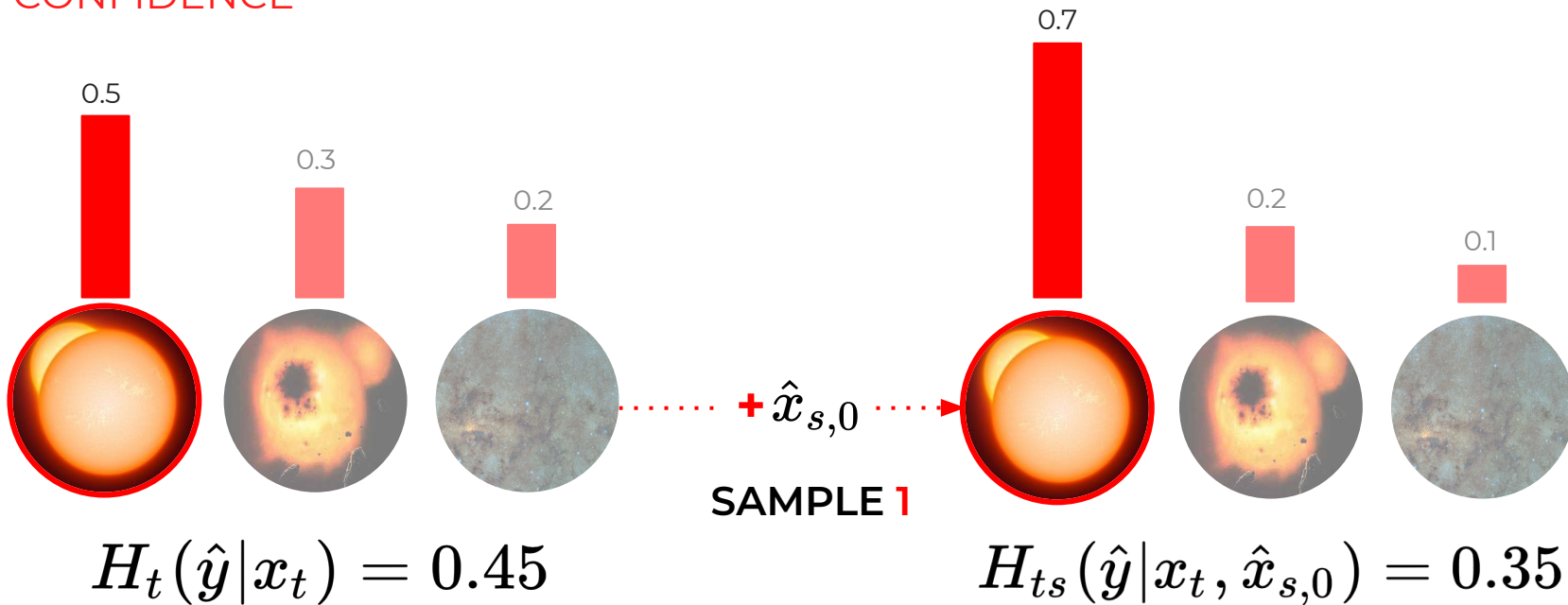
INTERESTING!

$$\Delta \hat{y}(x_t) = 1$$

CASE 2

CONFIDENCE

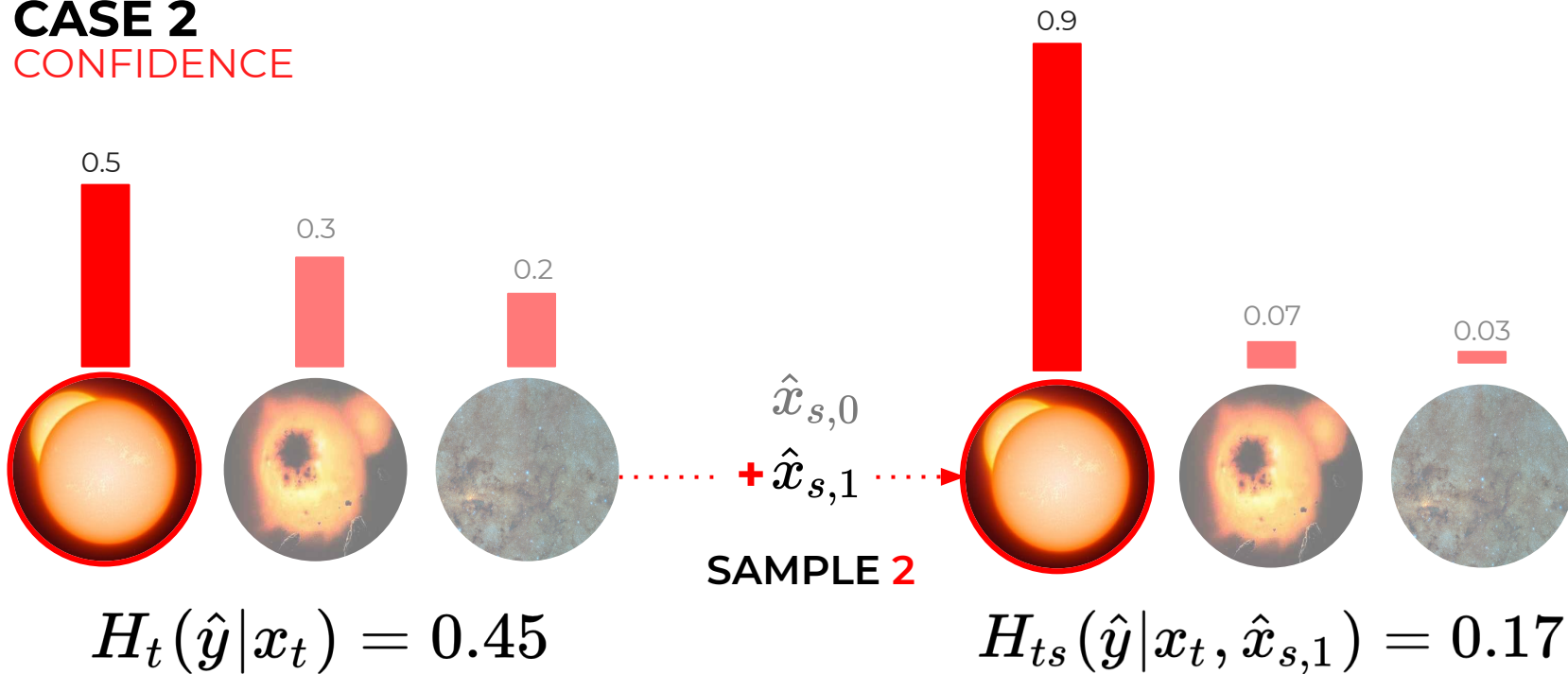
7



CASE 2

CONFIDENCE

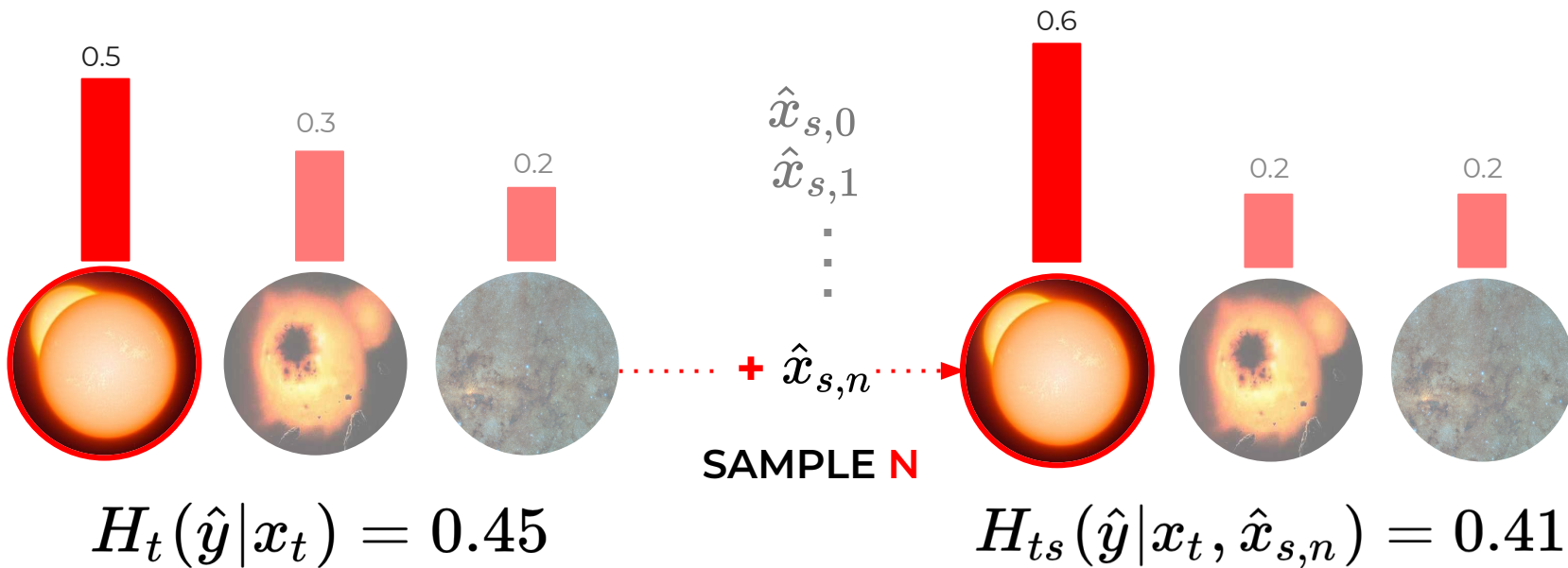
7



CASE 2

CONFIDENCE

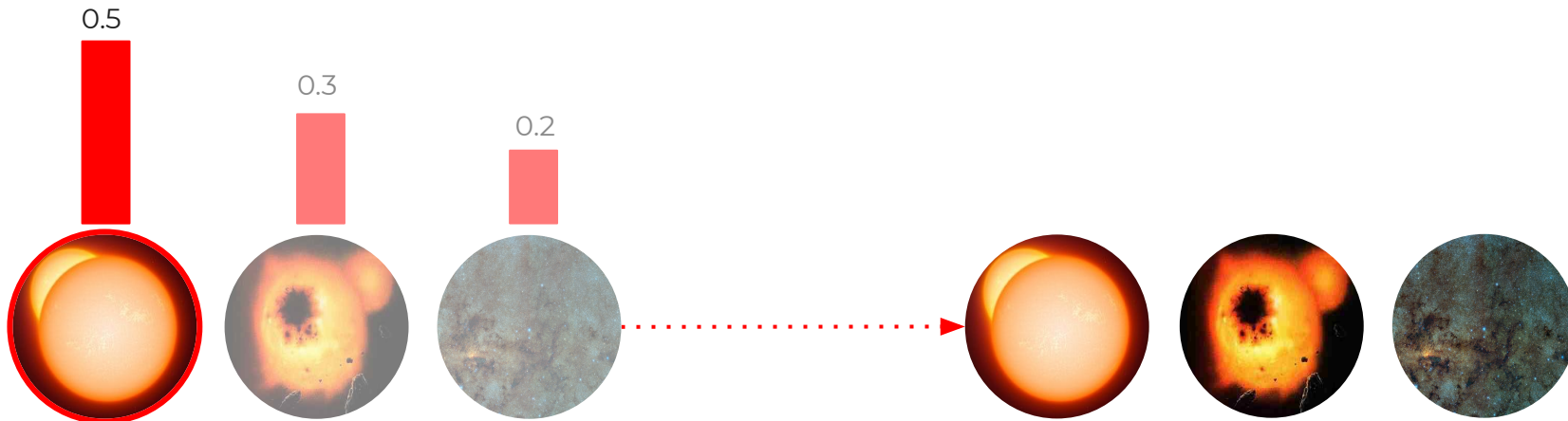
7



CASE 2

CONFIDENCE

7



$$H_t(\hat{y}|x_t) = 0.45$$

>

AVERAGE

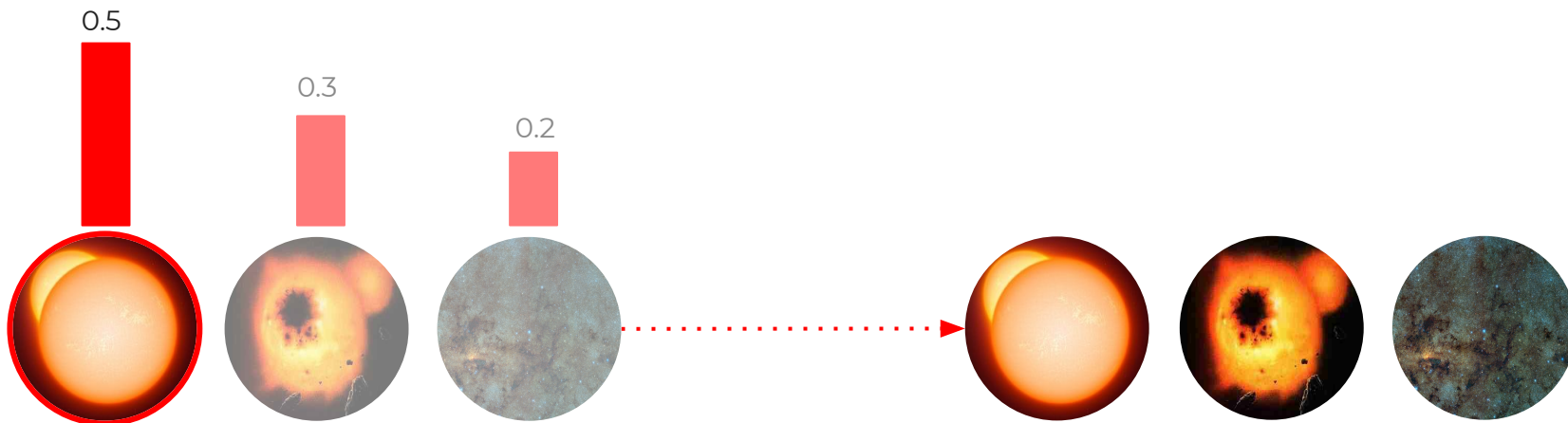
$$\overline{H}_{ts}(\hat{y}|x_t) = 0.31$$

LESS CONFUSED
INTERESTING!

CASE 2

CONFIDENCE

7



$$H_t(\hat{y}|x_t) = 0.45$$

>

$$\overline{H}_{ts}(\hat{y}|x_t) = 0.31$$

$$IG(x_t) = H_t(\hat{y}|x_t) - \overline{H}_{ts}(\hat{y}|x_t)$$

7

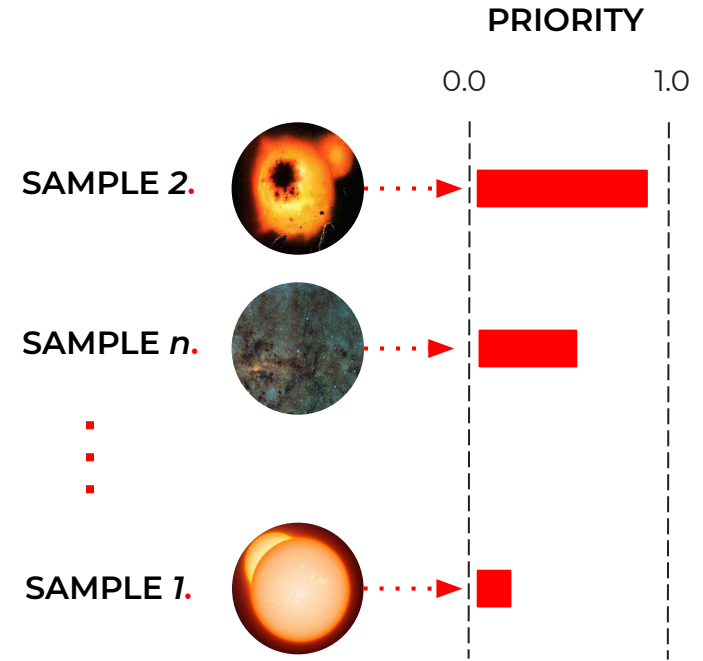
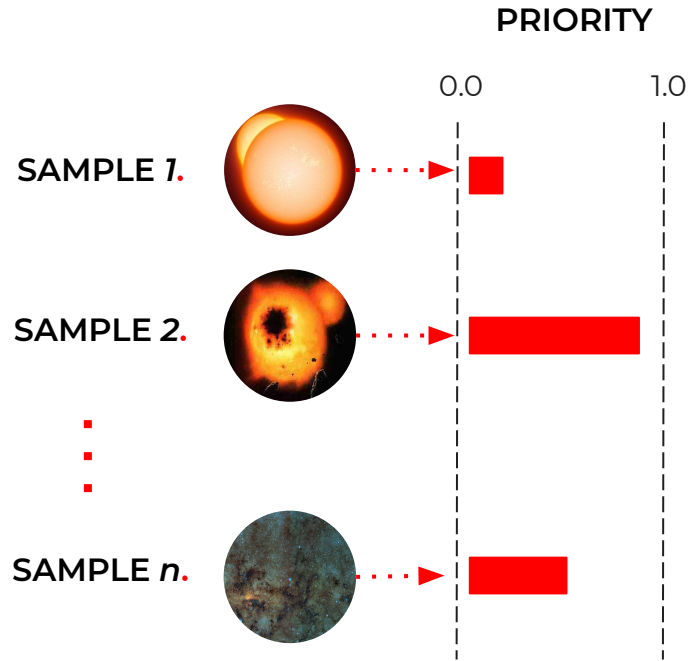
STRATEGIES

p : priority function

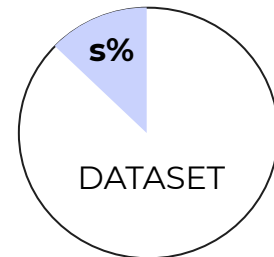
$$p : \{IG(x_t), \Delta \hat{y}(x_t)\} \longrightarrow [0, 1]$$

STRATEGIES

7



STRATEGIES

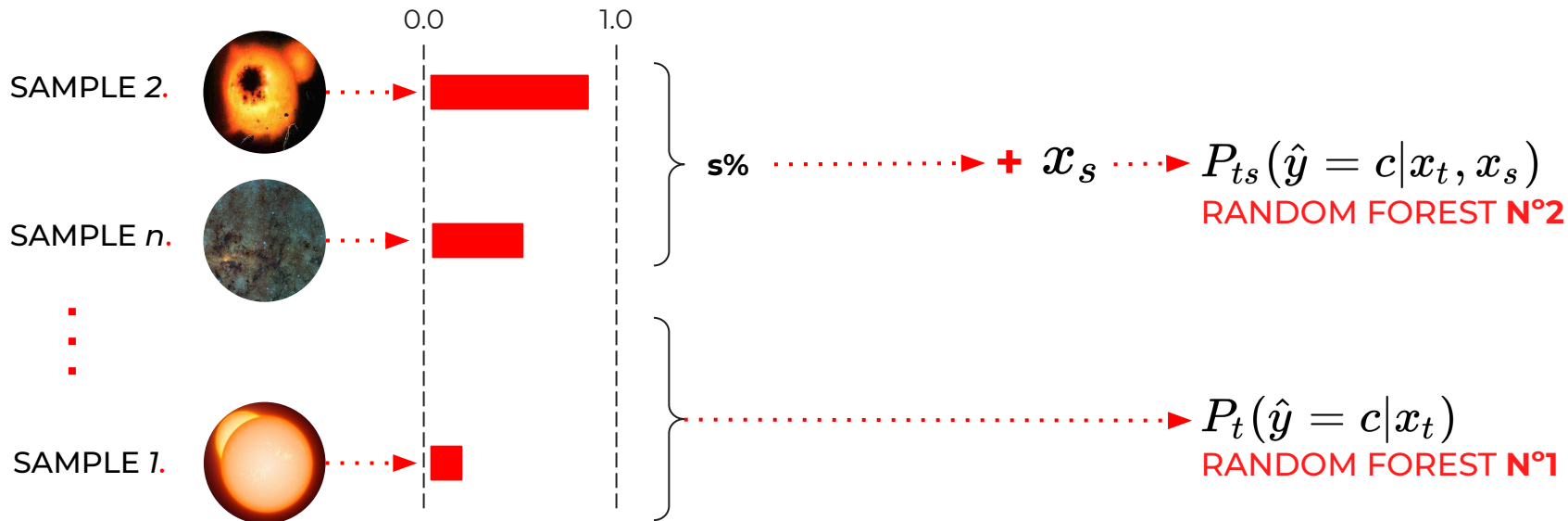


PRIORITY

ADD SPECTRUM

CLASSIFICATION

7



STRATEGIES

BASELINE



RANDOM



H_t

PROPOSED



IG



$IG + \Delta \hat{y}$



$H_t + \Delta \hat{y}$

FOR REFERENCE



IDEAL
SCENARIO

FEATURE EXTRACTION

FEATURE EXTRACTION

7

EXPERT FEATURES
FATS (Nun et al. 2015)

TIME SERIES

$$\vec{fats} = fats_1 \, fats_2 \, \dots \, fats_{44}$$

LEARNED FEATURES
VAE (Kingma et al. 2013)

SPECTRUM

$$\vec{\mu} = \mu_1 \, \mu_2 \, \dots \, \mu_7$$

OVERVIEW

OVERVIEW

1 TRAIN CLASSIFIERS

7

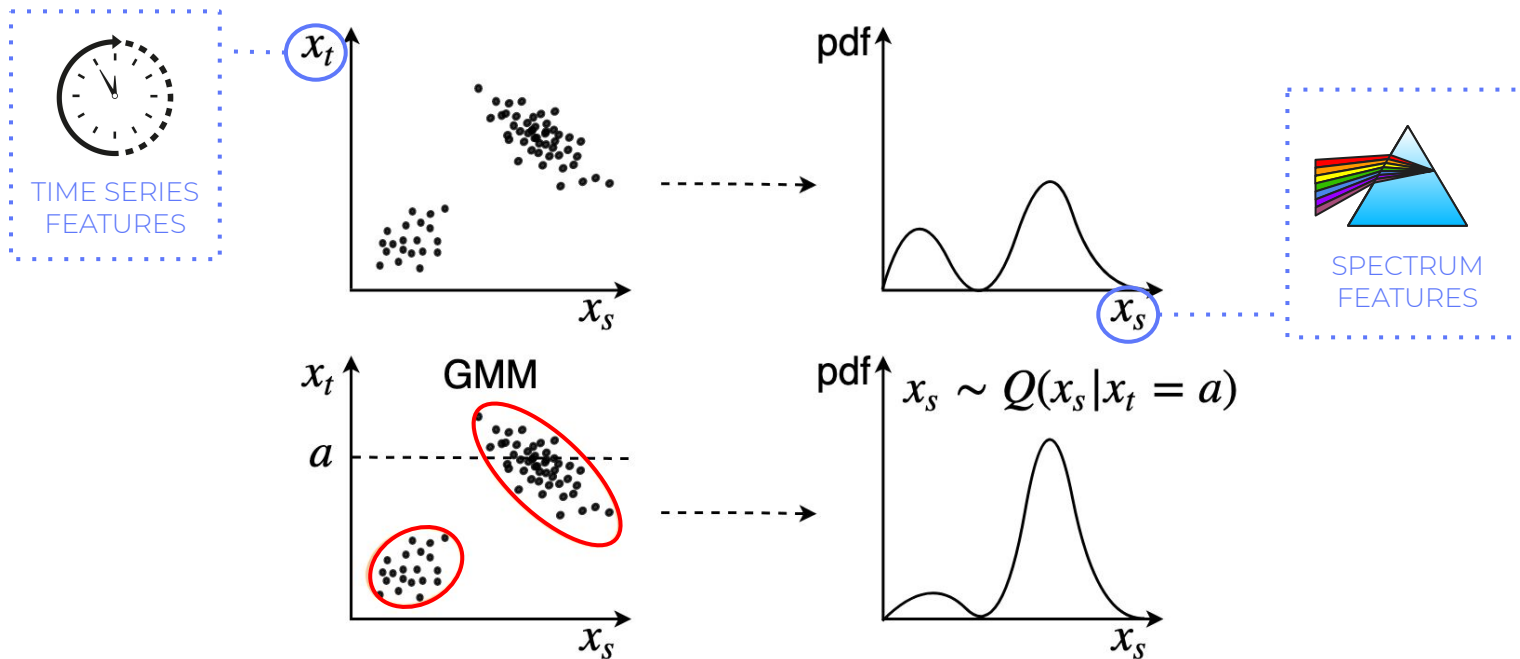
x_t $\rightarrow P_t(\hat{y} = c | x_t)$ RANDOM FOREST N°1

\neq

x_t, x_s $\rightarrow P_{ts}(\hat{y} = c | x_t, x_s)$ RANDOM FOREST N°2

OVERVIEW

2 MODEL FEATURE SPACE



OVERVIEW

3 METRICS APPROXIMATION

7

CASE 1

ACCURACY
Formerly wrongly
classified

$$\Delta \hat{y}(x_t, x_s) = 1 \quad \xrightarrow{\approx} \quad \Delta \hat{y}(x_t)$$

CASE 2

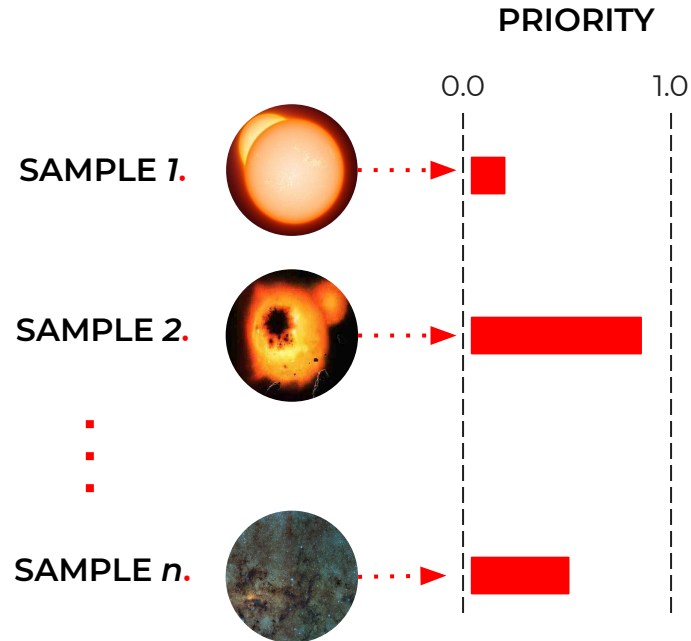
CONFIDENCE
Less confused
outcome

$$IG(x_t, x_s) > 0 \quad \xrightarrow{\approx} \quad IG(x_t)$$

OVERVIEW

4 STRATEGIES

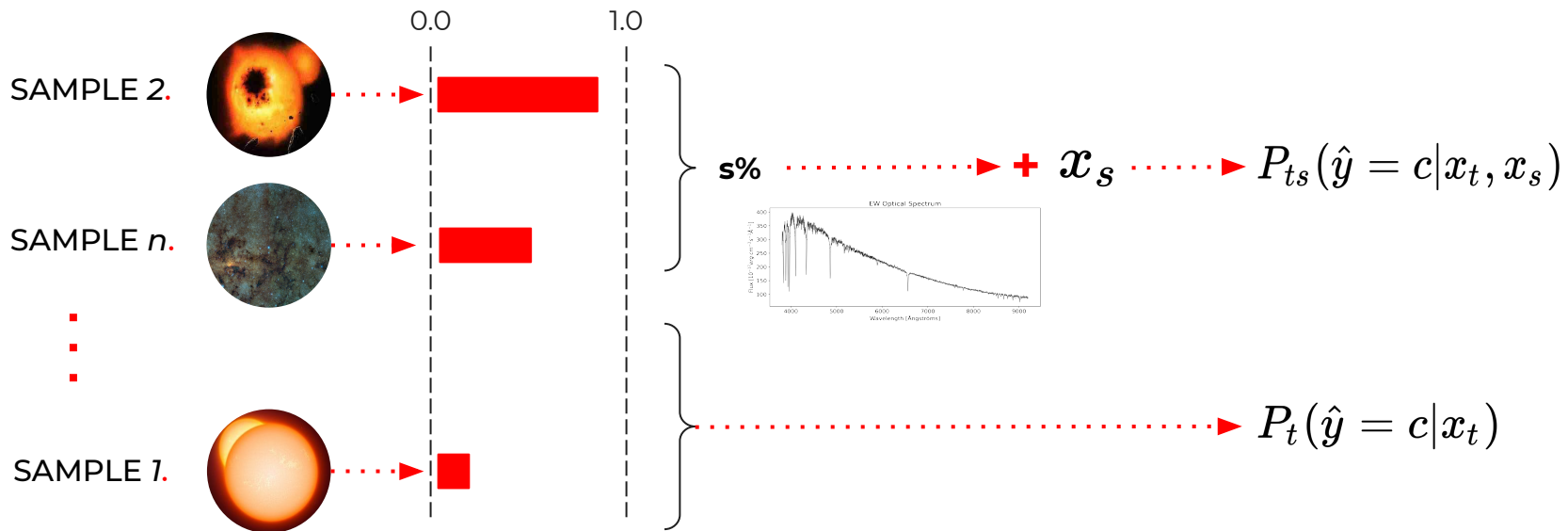
7



OVERVIEW

5 SELECTION AND EVALUATION

7



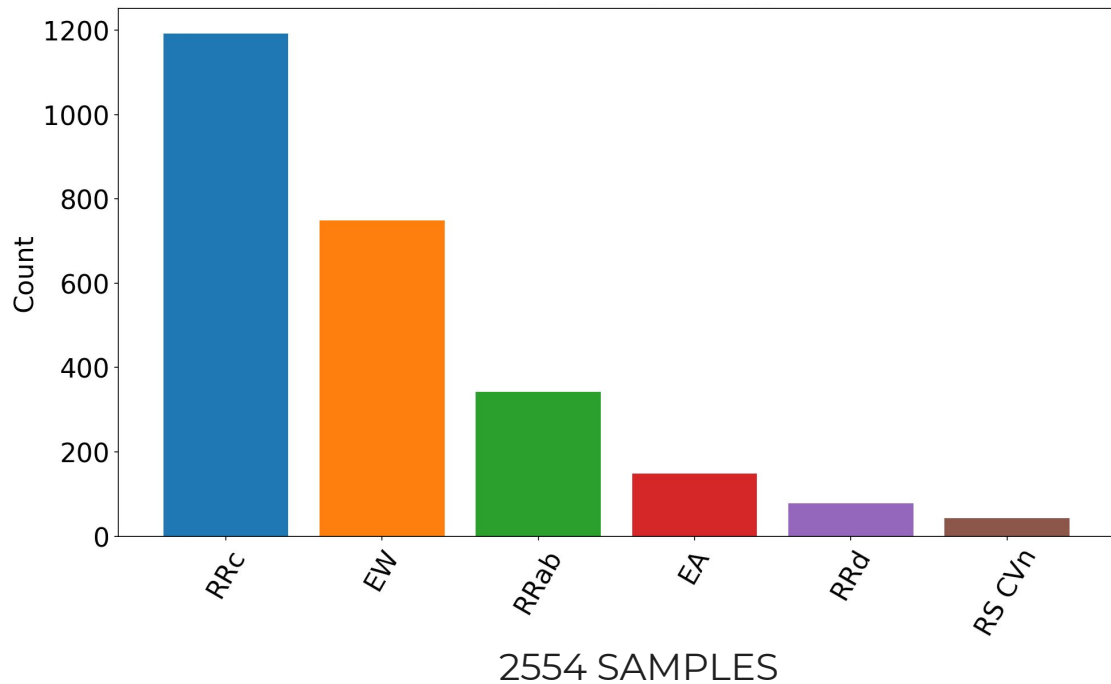
DATA

DATA.

7



Dataset Distribution



RESULTS

RESULTS

ACCURACY

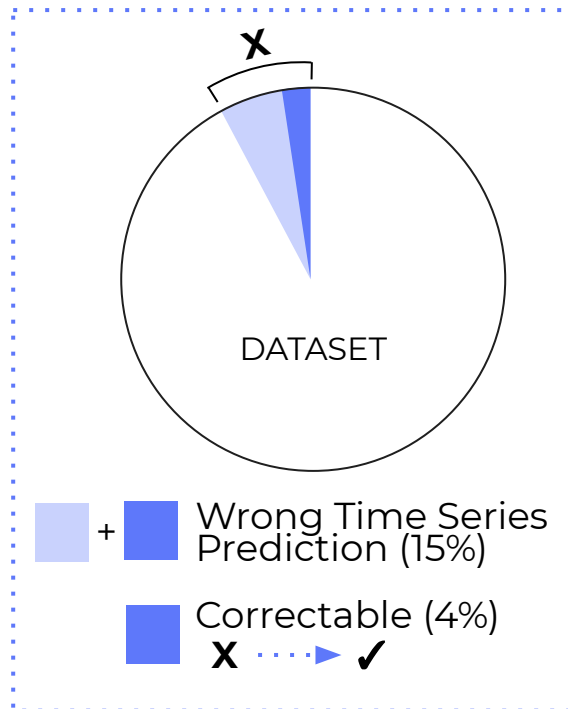
8

$$\Delta \hat{y}(x_t) = 1$$

113 POSITIVE OBJECTS = 4% DATASET

\$

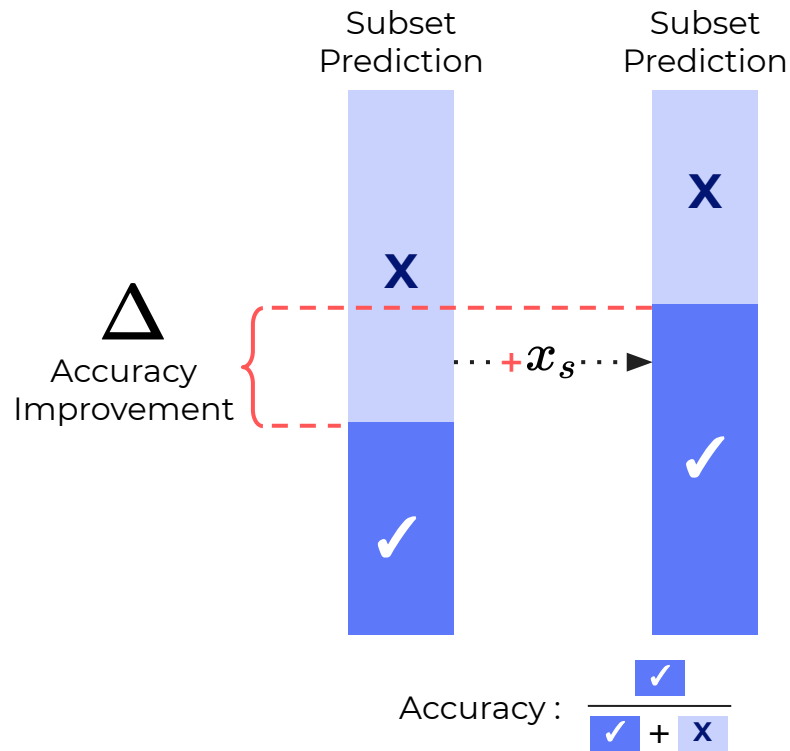
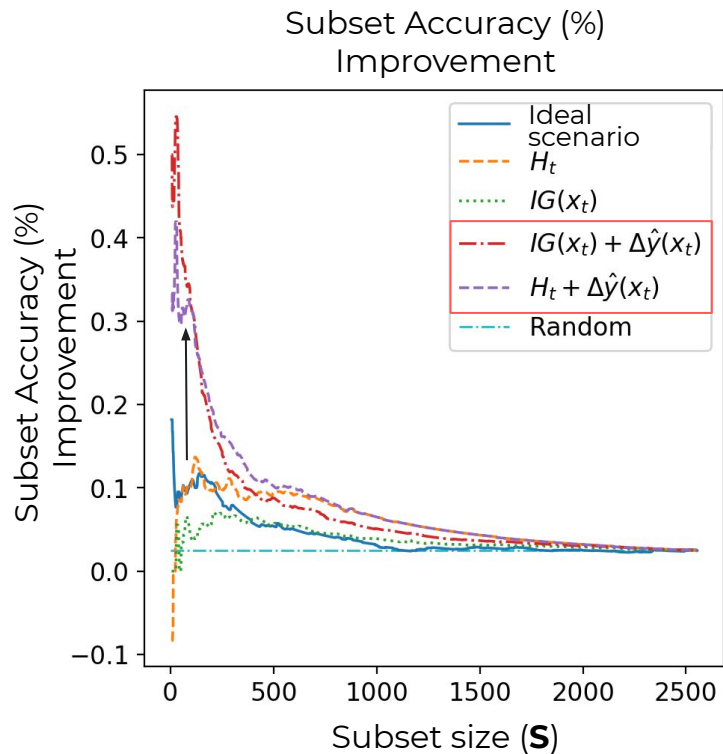
43 ARE CORRECTABLE = 46% OF AVAILABLE CORRECTABLE



RESULTS

ACCURACY

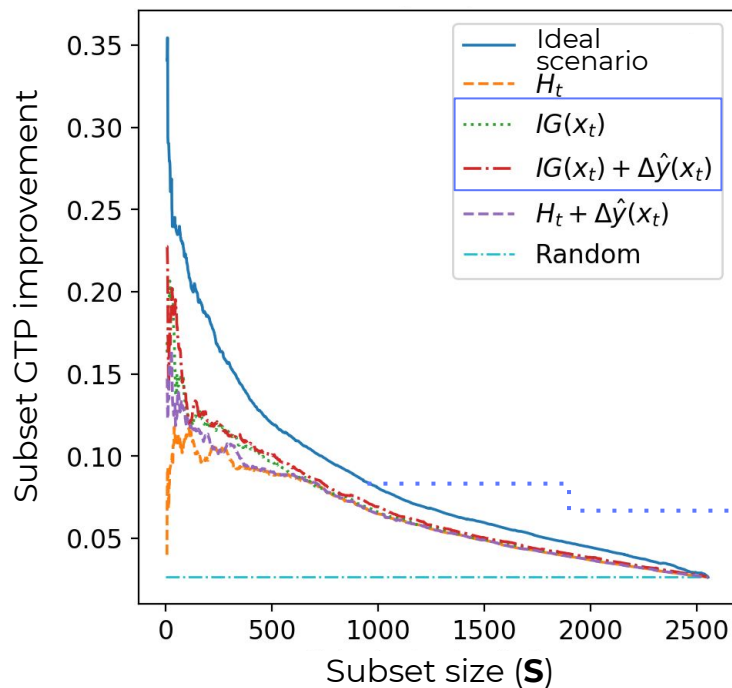
8



RESULTS

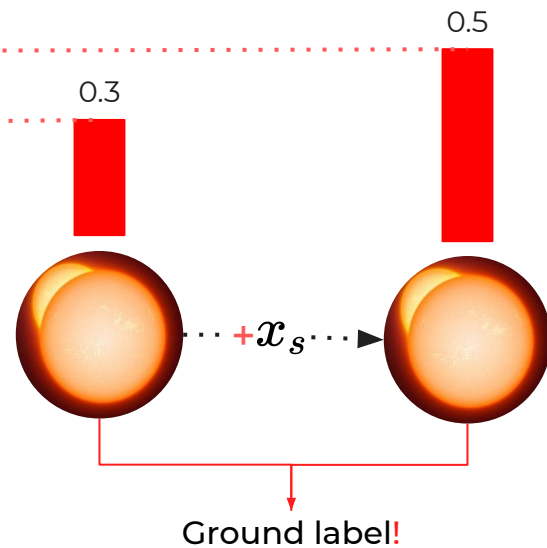
CONFIDENCE

GTP average improvement



Δ
GTP
Improvement

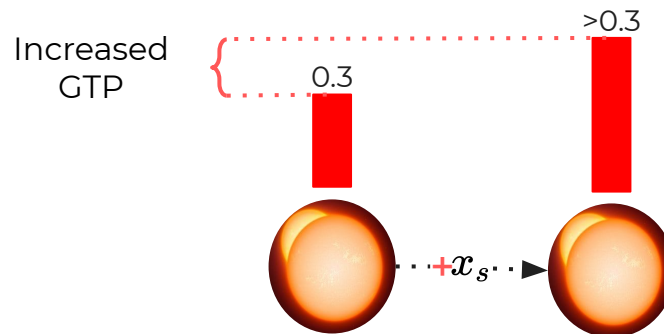
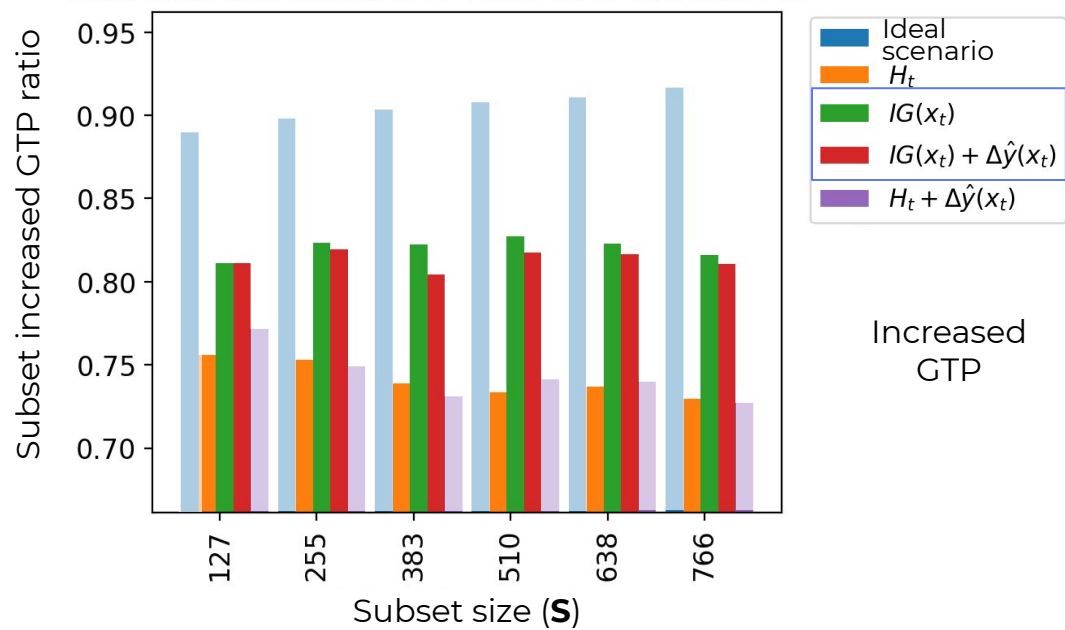
REFERENCE
As good as we
could do!



RESULTS

CONFIDENCE

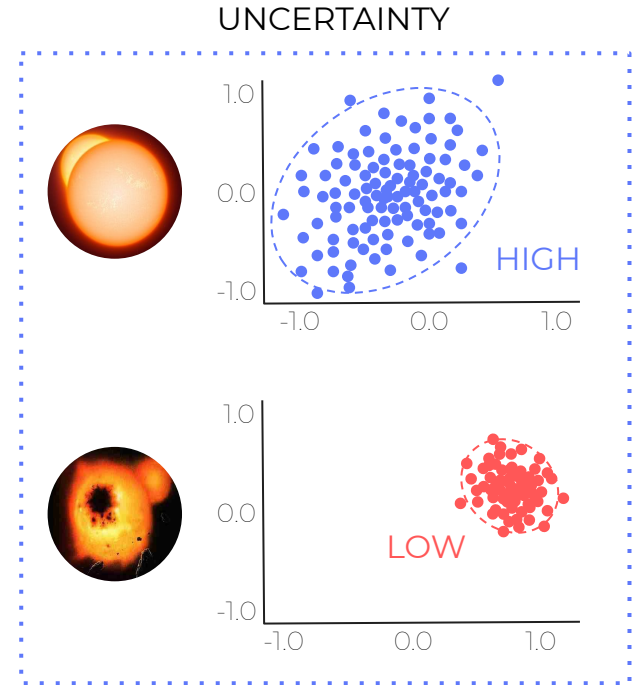
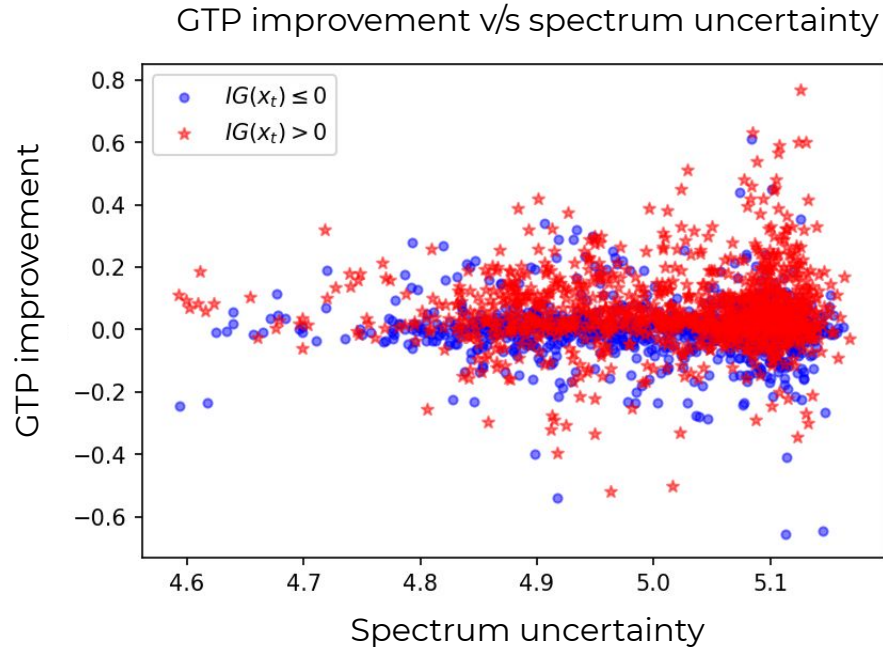
Objects with increased GTP ratio
over *Random* selection



RESULTS

CONFIDENCE

8



CONCLUSIONS AND FURTHER WORK

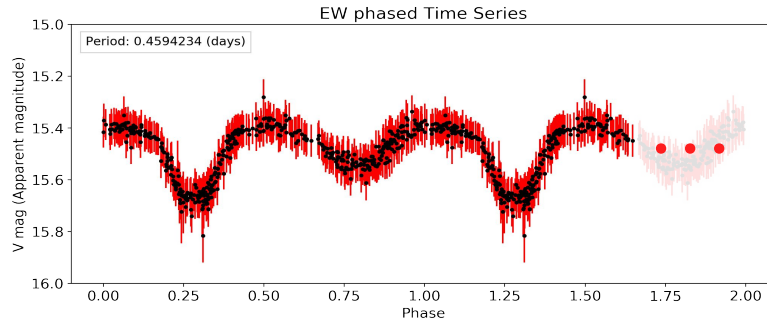
CONCLUSIONS

- $IG(x_t)$: SURPASSES BASE-LINE STRATEGIES Δ GTP
- $IG(x_t)$: IF FURTHER IMPROVED, MORE RESOURCES MAY BE SAVED
- $\Delta\hat{y}(x_t)$: DETECTION OF 46% OF CORRIGIBLE SAMPLES
- $\Delta\hat{y}(x_t)$: SURPASSES BASE-LINE STRATEGIES Δ ACCURACY
- WHICH STRATEGY? DEPENDS ON THE USE

FURTHER WORK

- TIME SERIES FEATURE LEARNING
- UNBALANCED CLASSES
- ONLINE SELECTION**
- $Q(x_t, x_s)$
- OPTIMUM NUMBER OF POINTS TO BE QUERIED
- OTHER METRICS

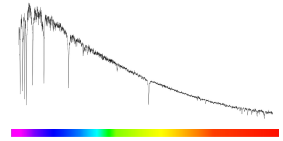
ONLINE SELECTION



OPTION 1
MORE EPOCHS



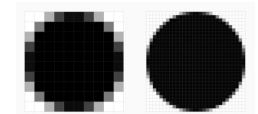
OPTION 3
SPECTRUM



OPTION 2
MORE COLOR
BANDS



OPTION 4
SPECTRUM
RESOLUTION



QUESTIONS?

RESULTS

ACCURACY

Subset correctable ratio over *Random* selection

