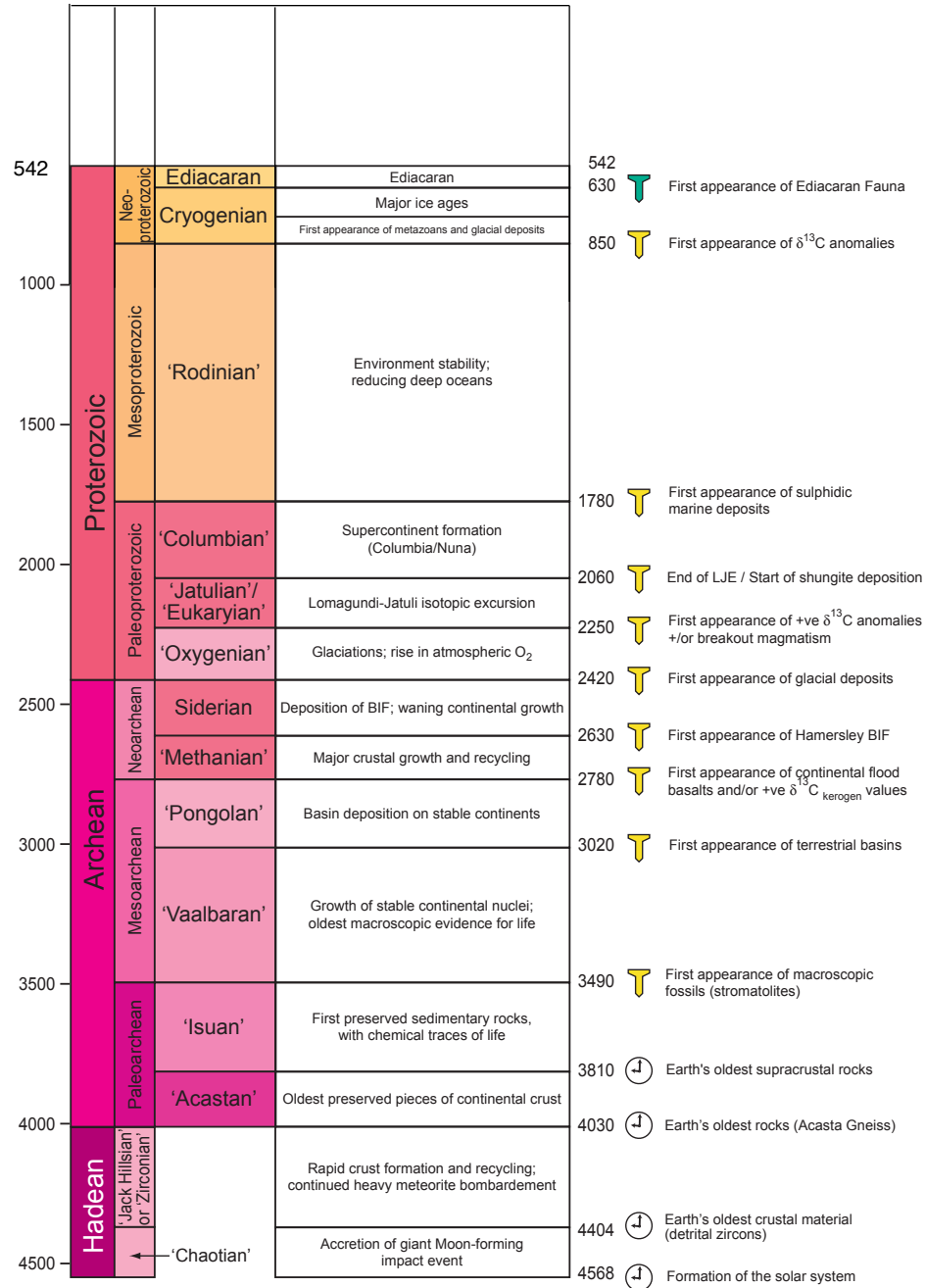
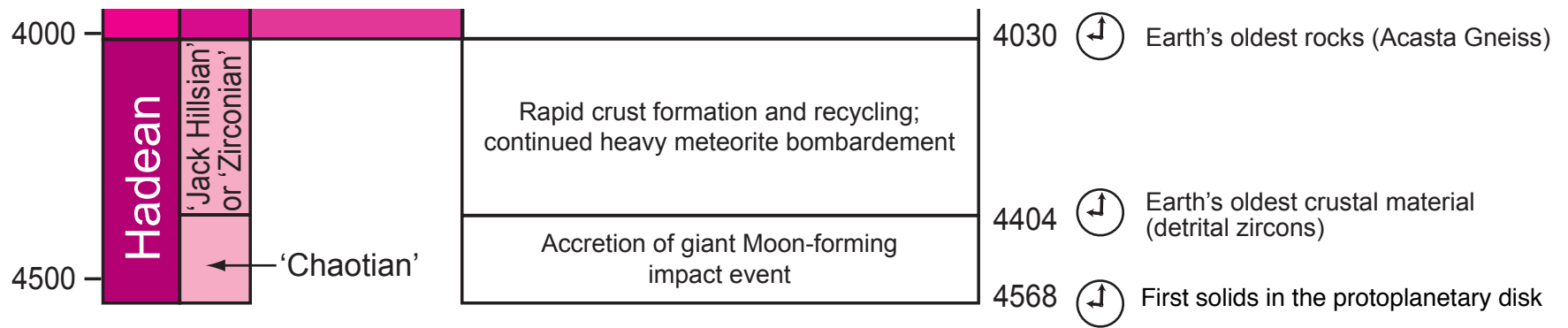
An aerial photograph of a tropical coastline, showing a large bay or inlet. The water is a deep blue-grey, and the surrounding land is covered in dense, dark green vegetation, likely palm trees. The sky is overcast and grey, creating a moody atmosphere. The text is overlaid on the upper half of the image.

The real 1%:
Volatiles in planetary accretion and the
rapid development of habitability

Lindy Elkins-Tanton
DTM, Carnegie Institution

What period of time are we talking about?





How do rocky planets obtain atmospheres?

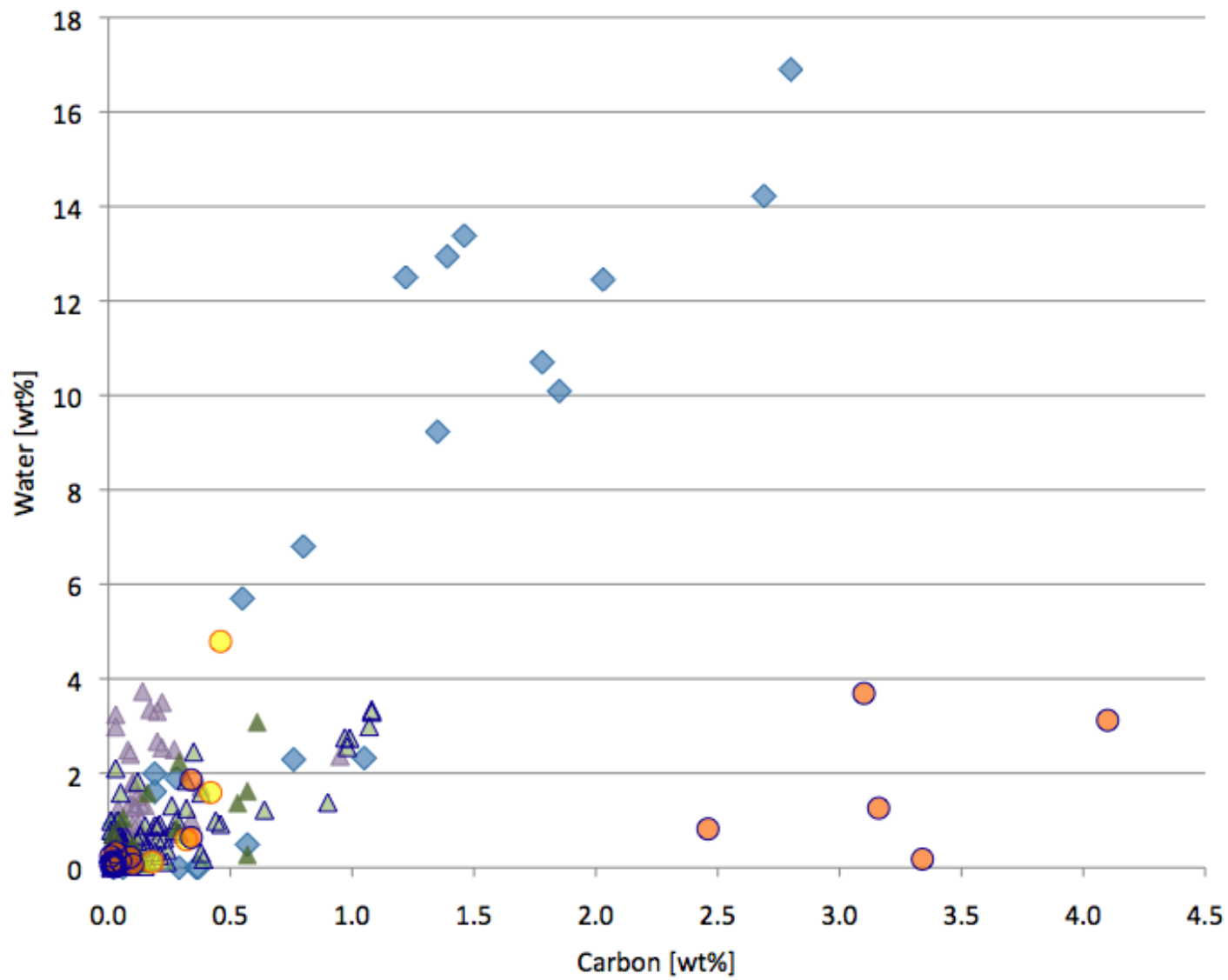
1. Capture of nebular gases
2. Degassing from the interior
3. Later surface impacts

What material are we working with?

TABLE 5
PHOTOSPHERIC Z/X IN ELEMENTAL
ABUNDANCE COMPILATIONS

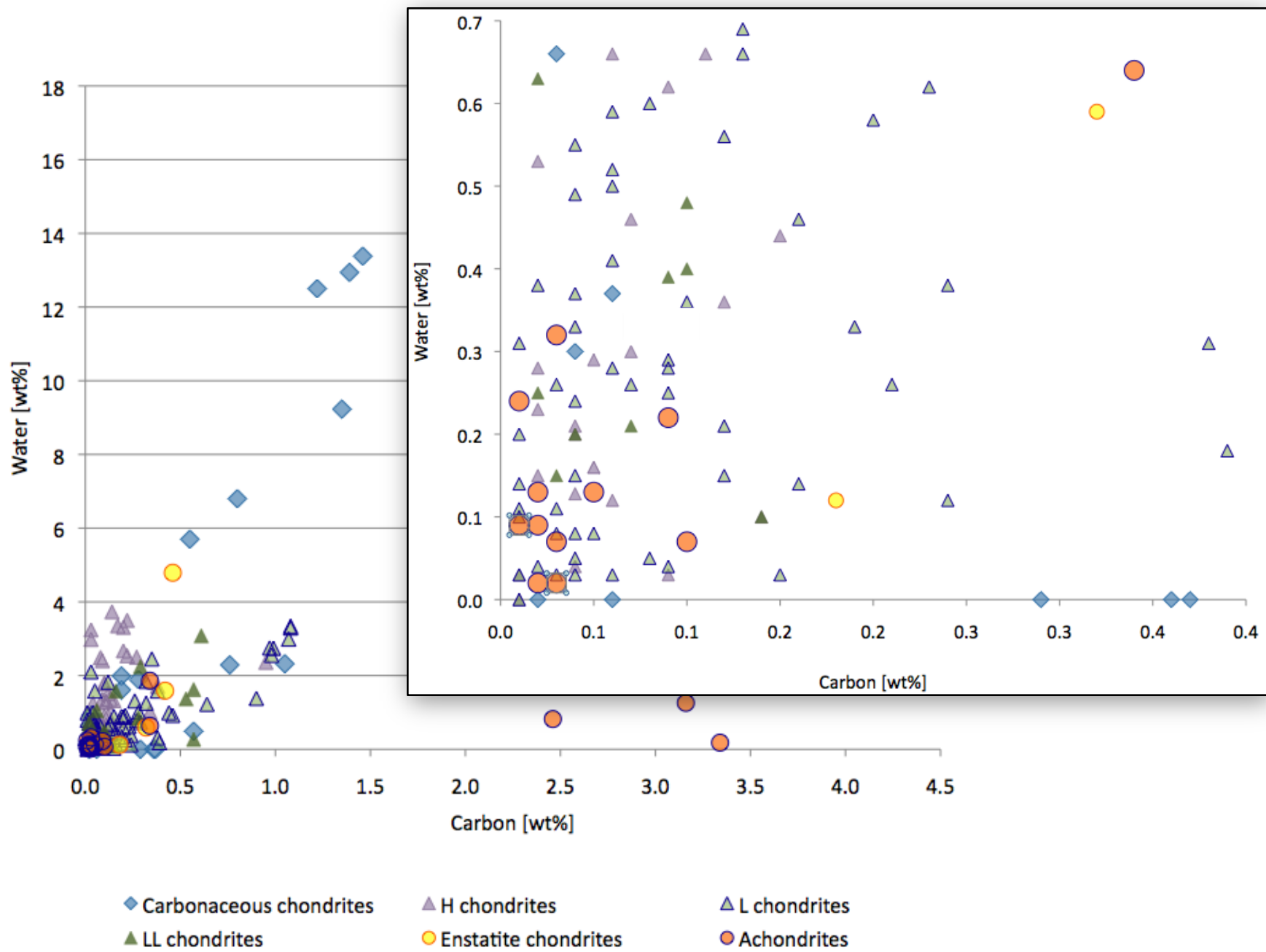
Z/X	Year	Reference
0.0270	1984	1
0.0267	1989	2
0.0245	1993	3
0.0244	1996	4
0.0229	1998	5
0.0208	2002	6
0.0177	2003	7

REFERENCES—(1) Grevesse 1984; (2) Anders & Grevesse 1989; (3) Grevesse & Noels 1993; (4) Grevesse et al. 1996; (5) Grevesse & Sauval 1998; (6) Grevesse & Sauval 2002; (7) this work.



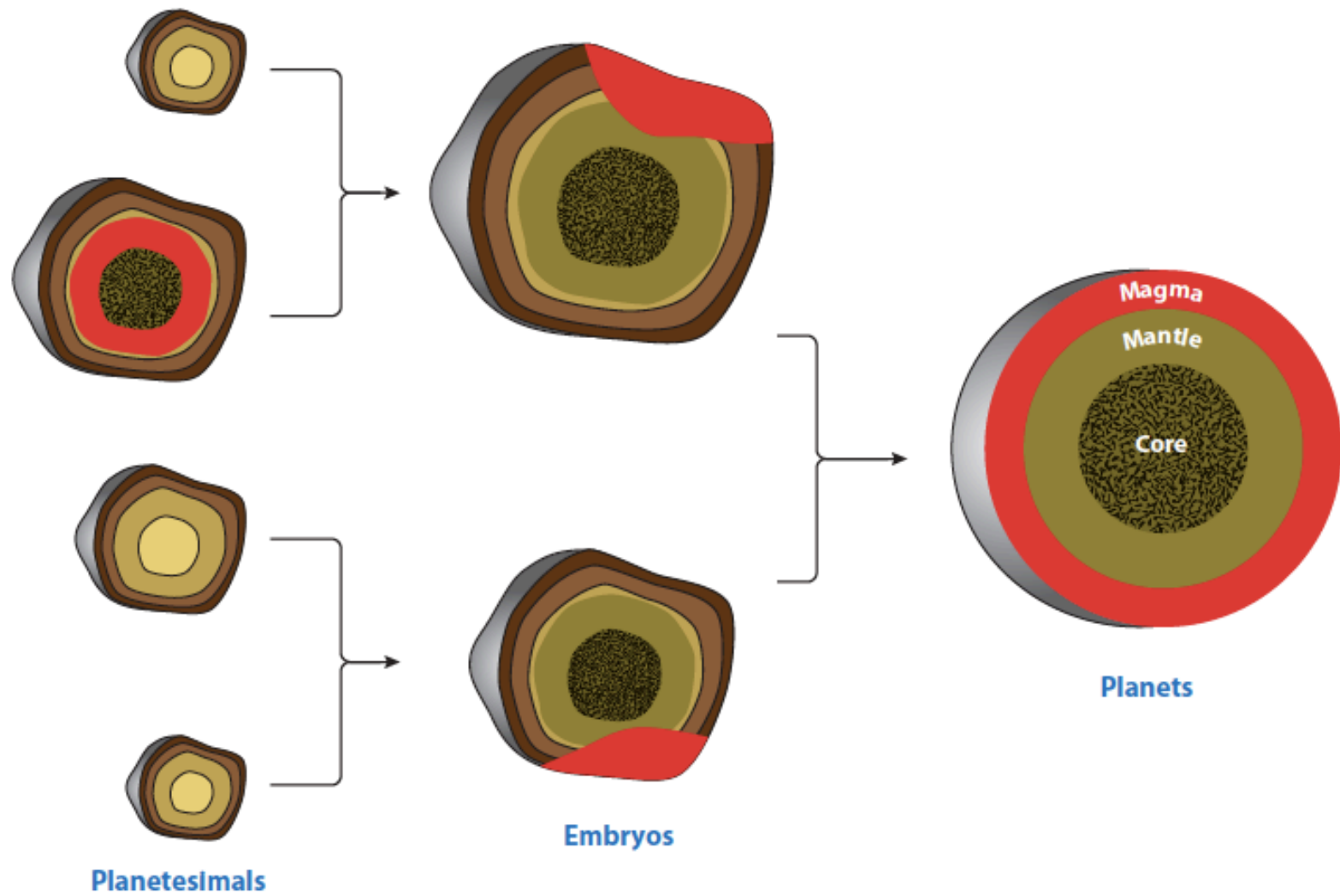
- ◆ Carbonaceous chondrites
- ▲ H chondrites
- ▲ L chondrites
- ▲ LL chondrites
- Enstatite chondrites
- Achondrites

Data from Jarosewich (1990)

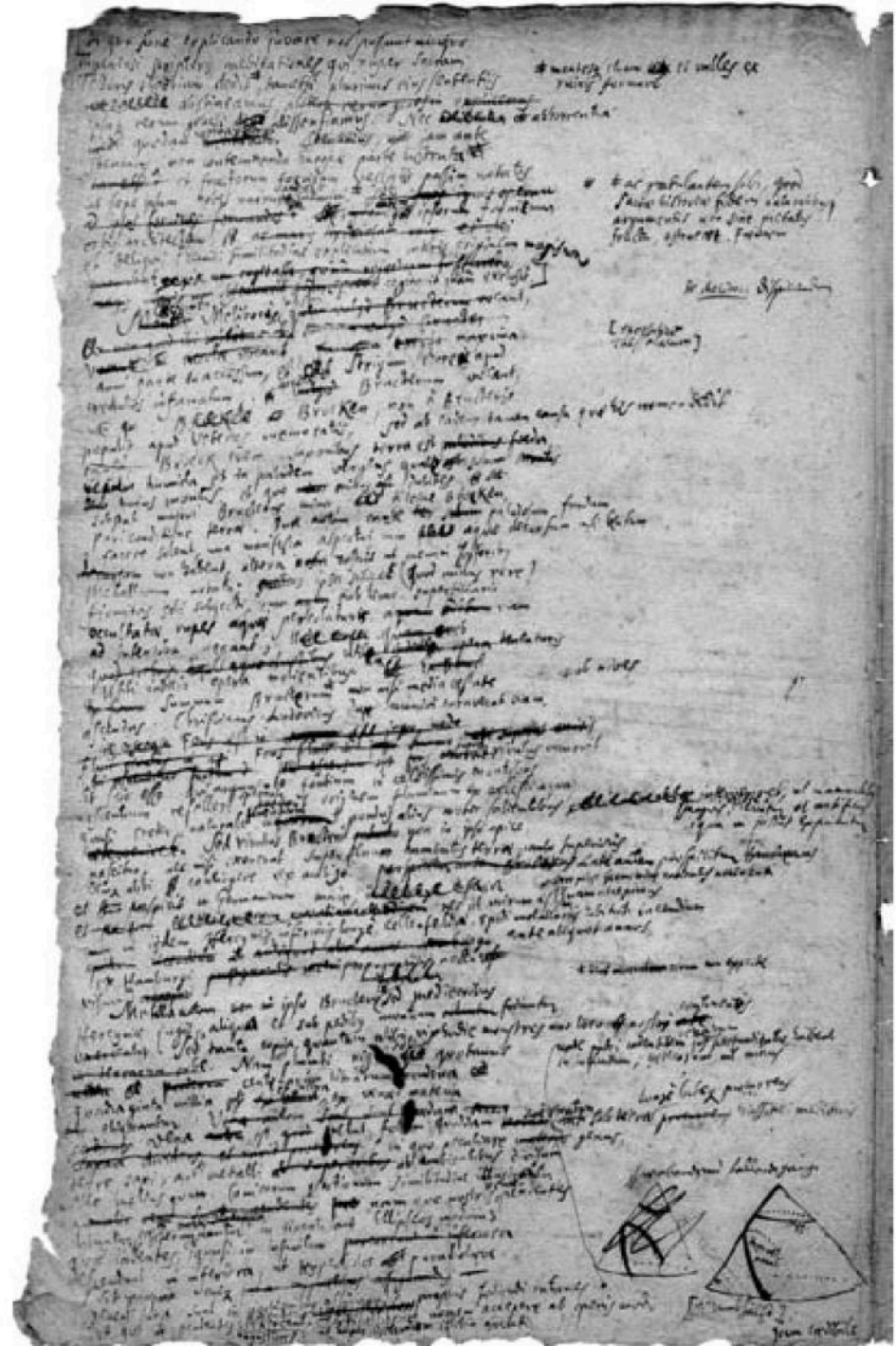


Data from Jarosewich (1990)

What processes make planets?

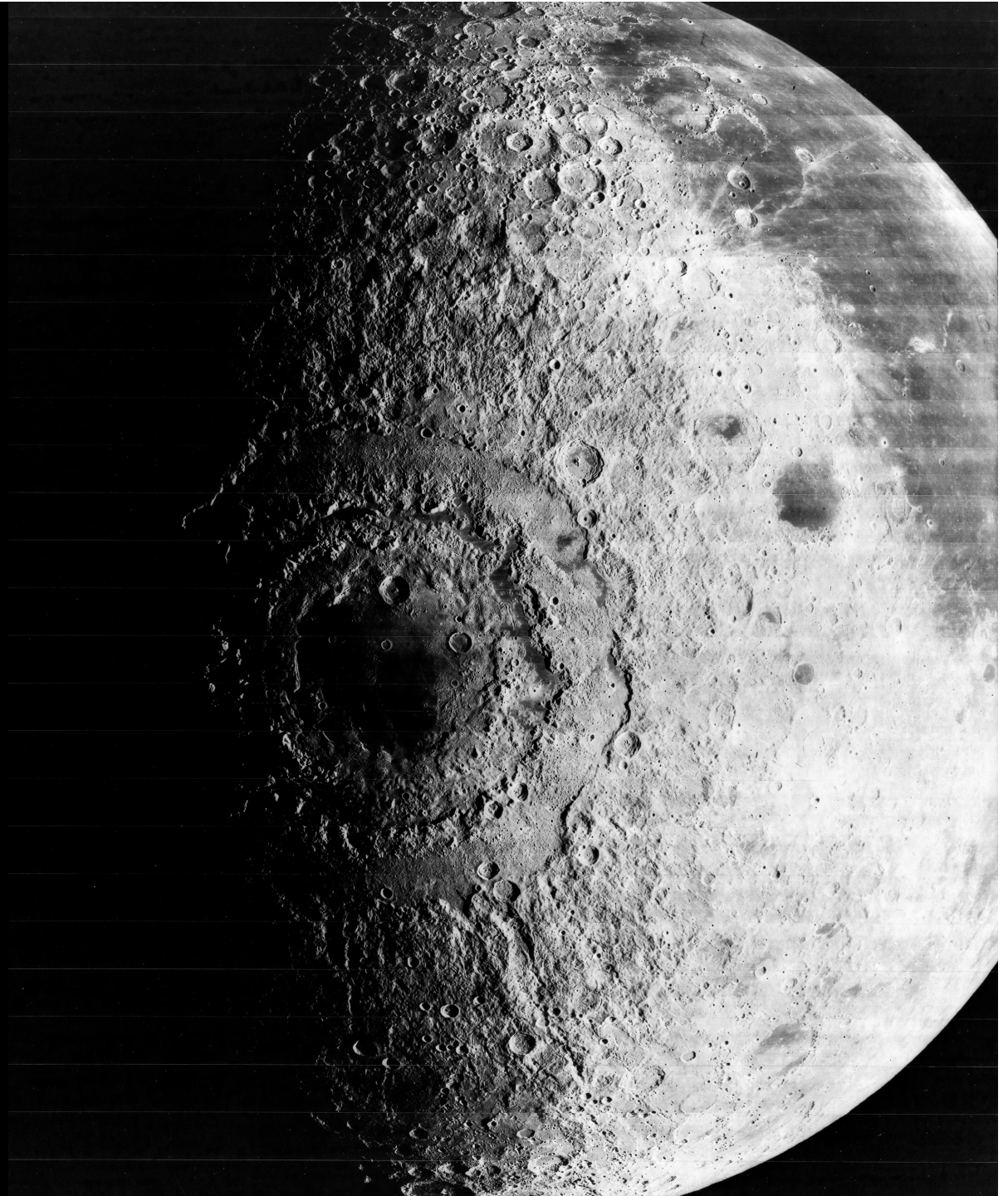


1693: Leibniz suggests in his book *Protogaea* that the Earth was once molten



Planets retain volatiles during accretion

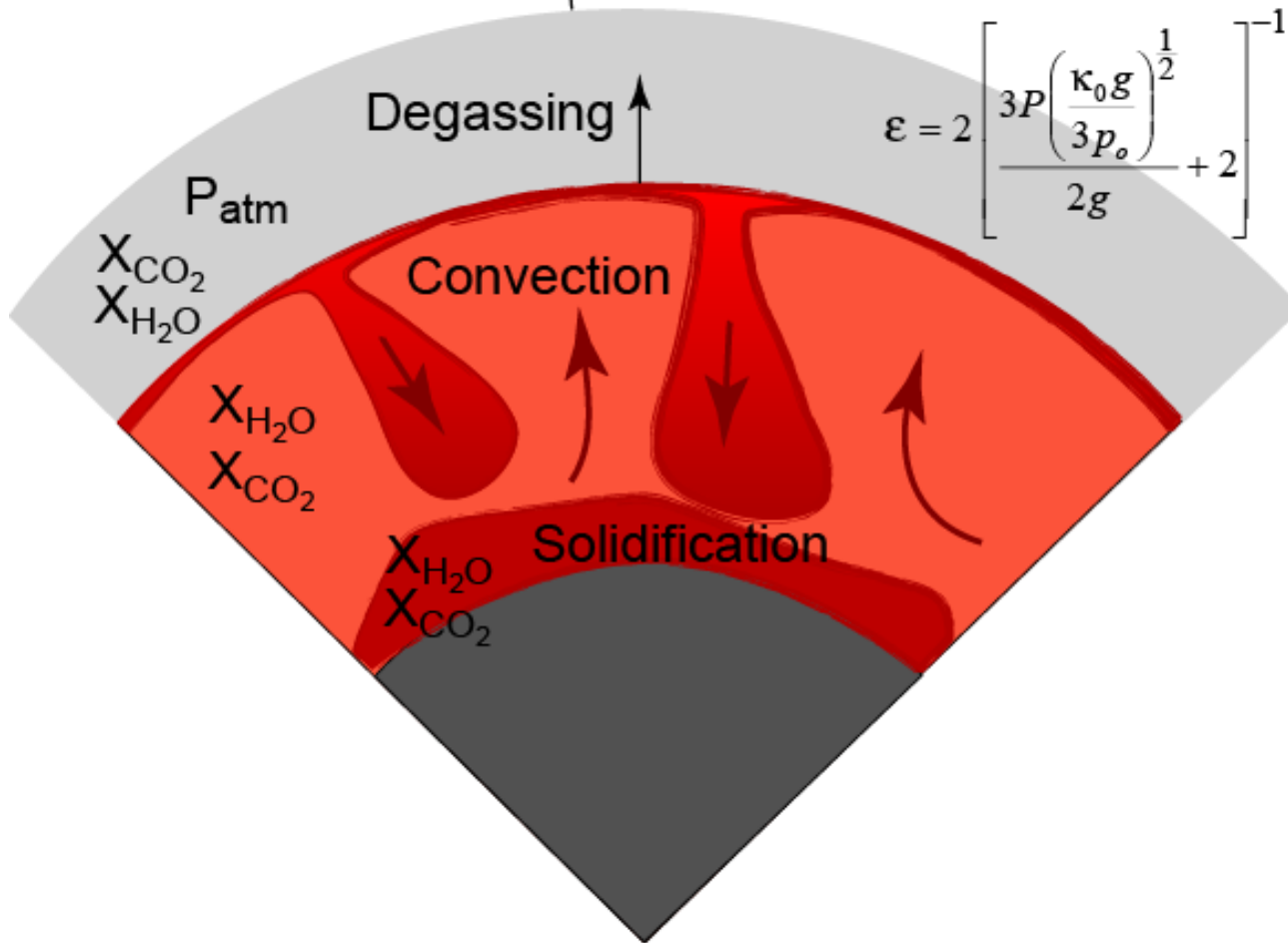
- Moon, Mars, Mercury
- Water, carbon, ..., helium!



$$4\pi R^2 F = V \left[\rho H 4\pi r^2 + \rho C_p \frac{dT}{dr} \frac{4}{3} \pi (R^3 - r^3) \right]$$

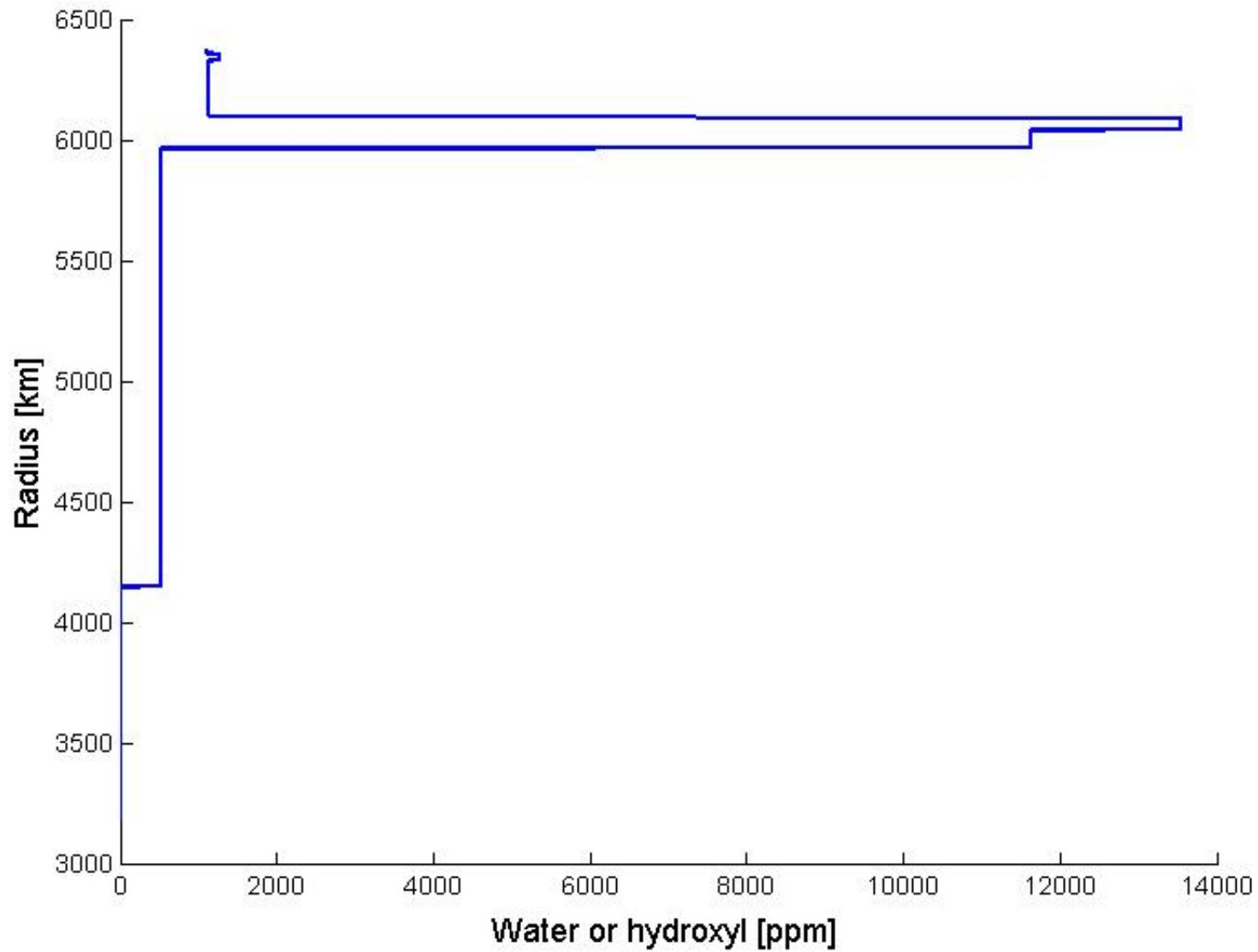
Heat flux \uparrow $F = \epsilon \sigma (T^4 - T_\infty^4)$

$$\epsilon = 2 \left[\frac{3P \left(\frac{\kappa_0 g}{3p_0} \right)^{\frac{1}{2}}}{2g} + 2 \right]^{-1}$$

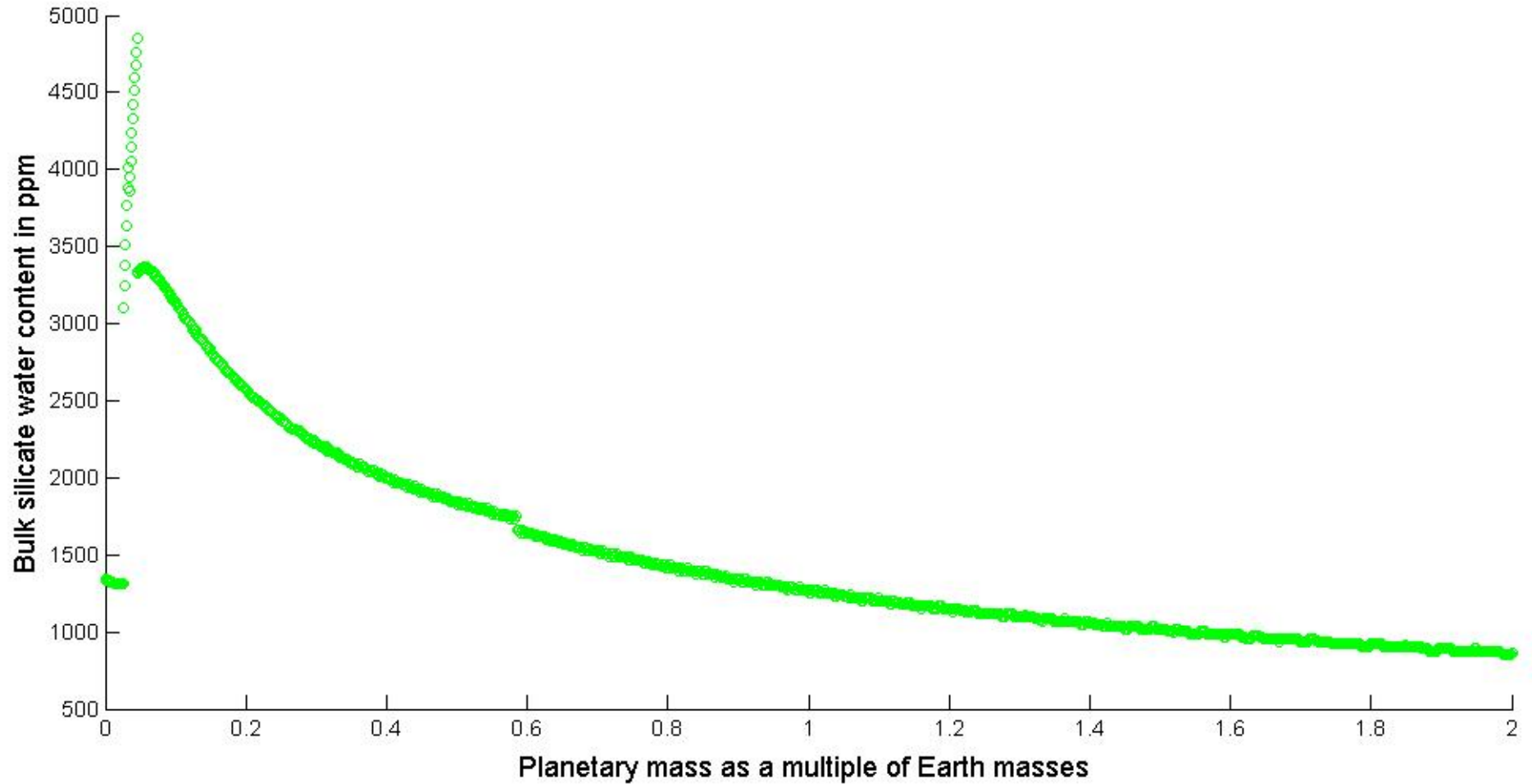


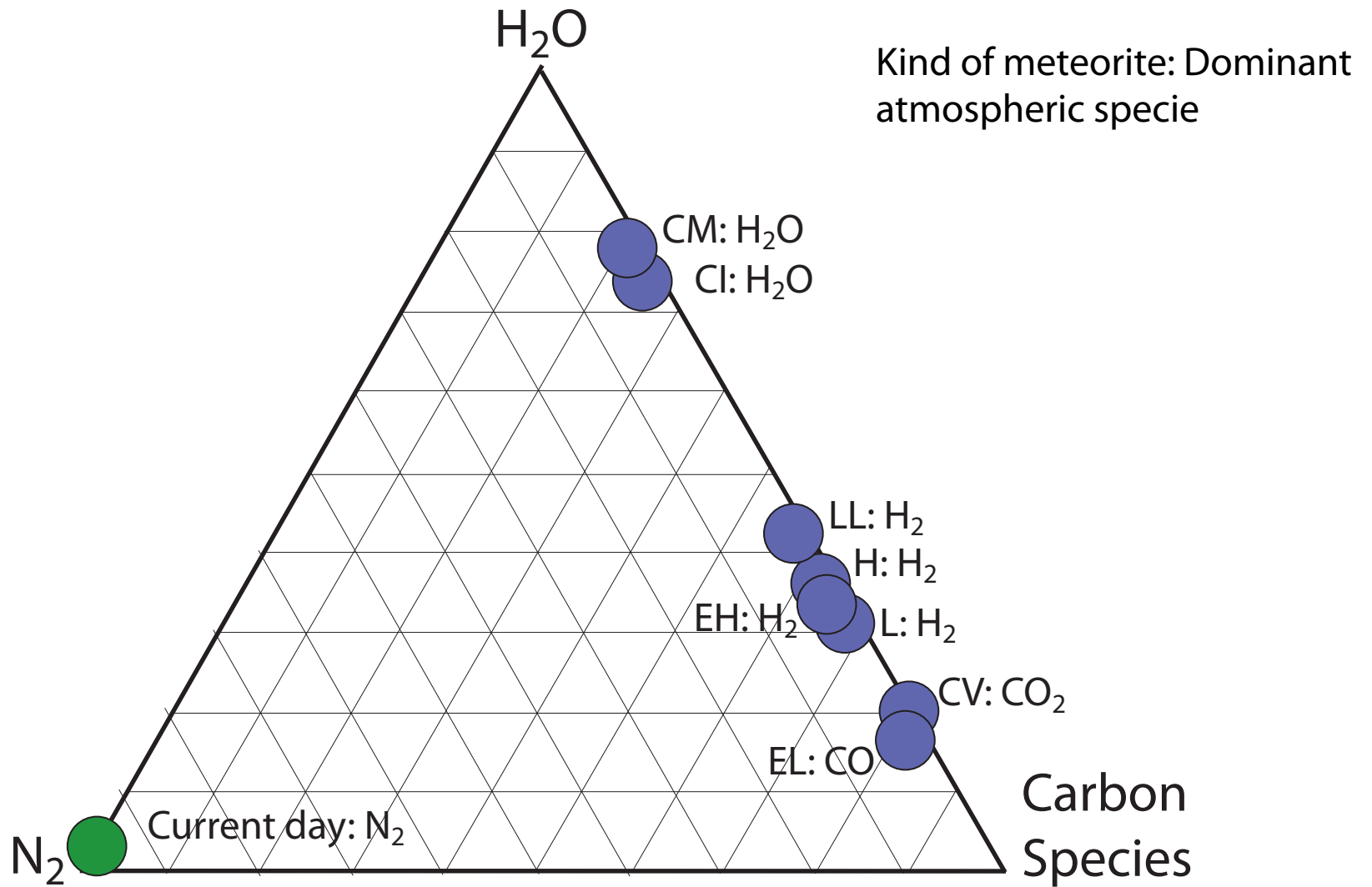
Some volatiles inside, some volatiles outside.

The maximum water capacities of mantle minerals vary widely



Maximum bulk water content retained in planetary mantle





Earliest degassed atmospheres vary with oxidation state, and all are a long way from current day

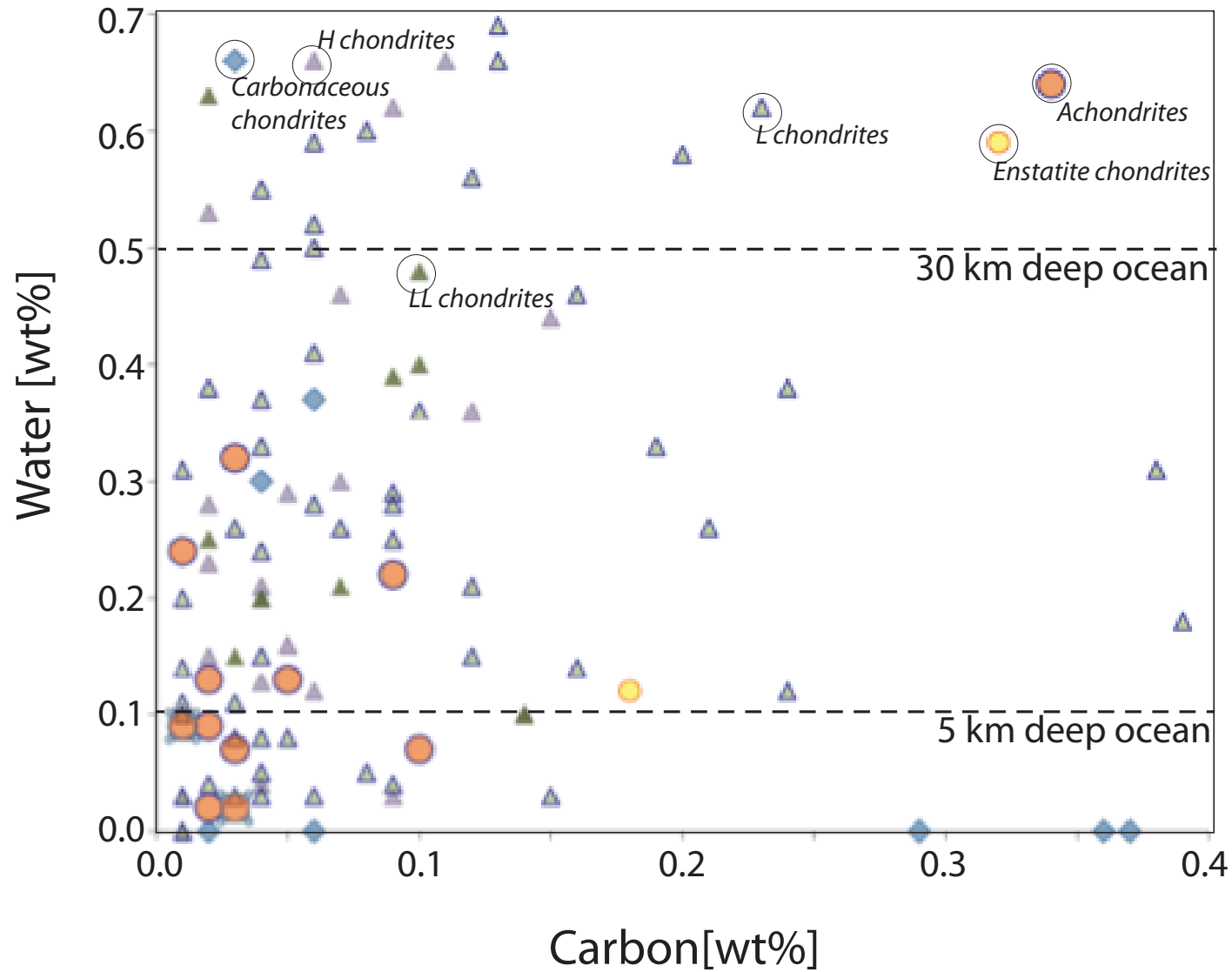
DATA from Schaefer and Fegley (2010), Hashimoto et al. (2007), Elkins-Tanton and Seager (2008)



Life (1952)



New York Times (2008)



Earth starting with 100 ppm bulk water produces dense atm that collapses upon cooling into global ocean 100s m deep

Elkins-Tanton ; Data from Jarosewich (1990)

